

TUNNELING IN QUANTUM FIELD THEORY

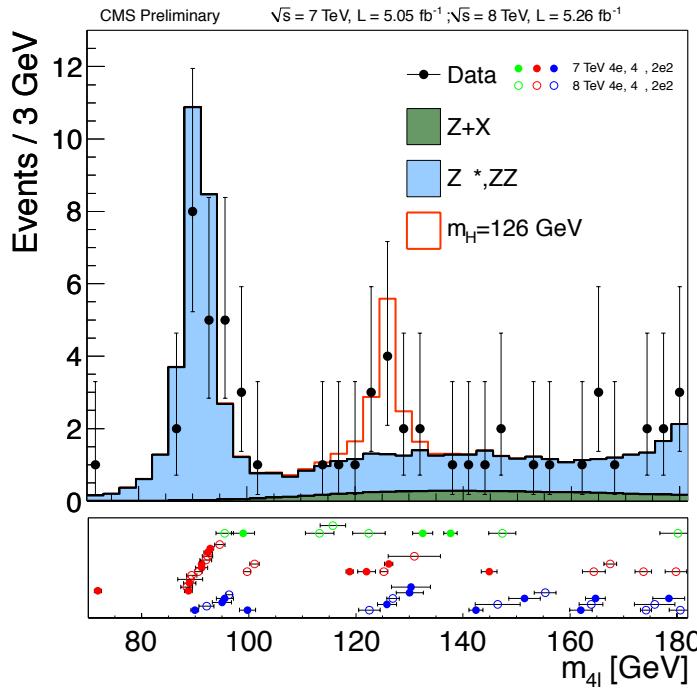
SLAC

November 18, 2016

Matthew Schwartz
Harvard University

Based on PRL 113.241801 (arXiv:1408.0292)
PRD (91) 016009 (arXiv:1408.0287)
PRL 118.xxxx [to appear] (arXiv:1602.01102)
PRD [in review] (arXiv:1604.06090)
with Anders Andreassen, David Farhi and William Frost

July 4, 2012: Higgs boson discovered!



What did we learn?

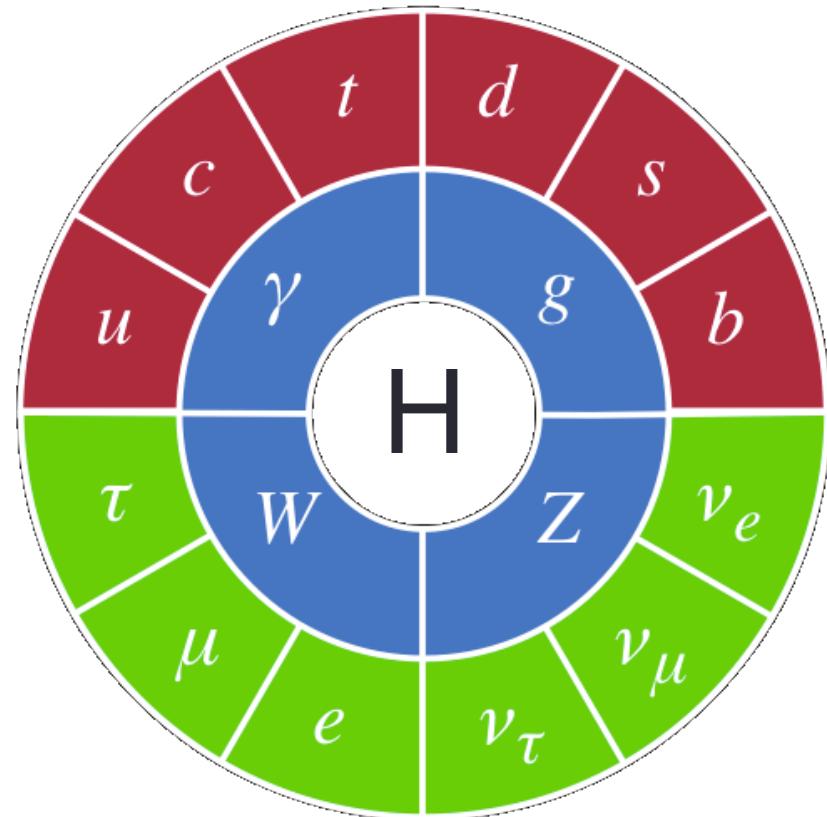
The Standard Model

1980-2012

u up	s strange	t top	γ photon
d down	c charmed	b bottom	g gluon
ν_e electron neutrino	ν_μ muon neutrino	ν_τ tau neutrino	Z Z Boson
e electron	μ muon	τ tau	W W Boson

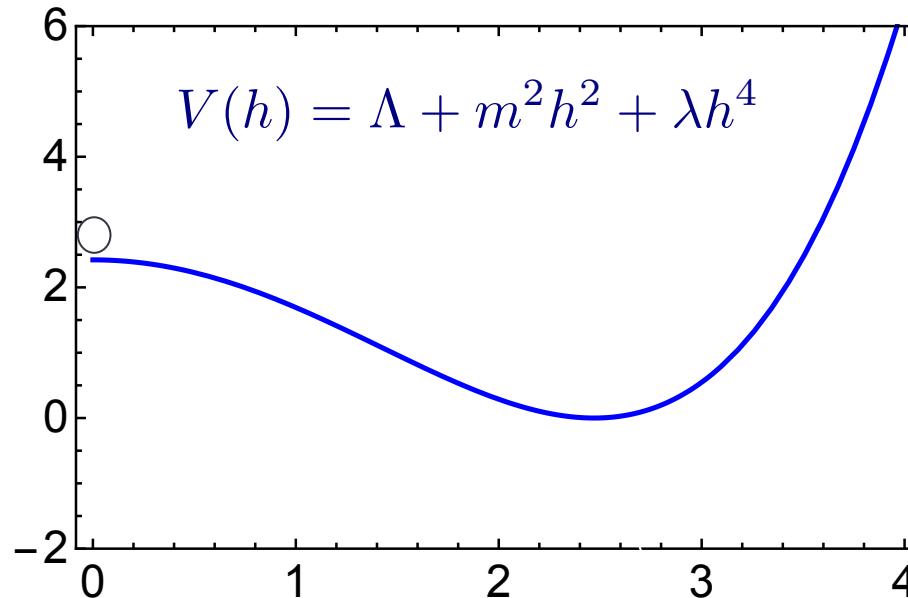


2012 -- ??



What is the Higgs field?

- The Higgs field $h(x)$ pervades all space
- The Higgs field $h(x)$ has charge under the weak force
 - If $\langle h \rangle = 0$ space is not empty – it has weak charge too
- The Higgs field $h(x)$ has a potential



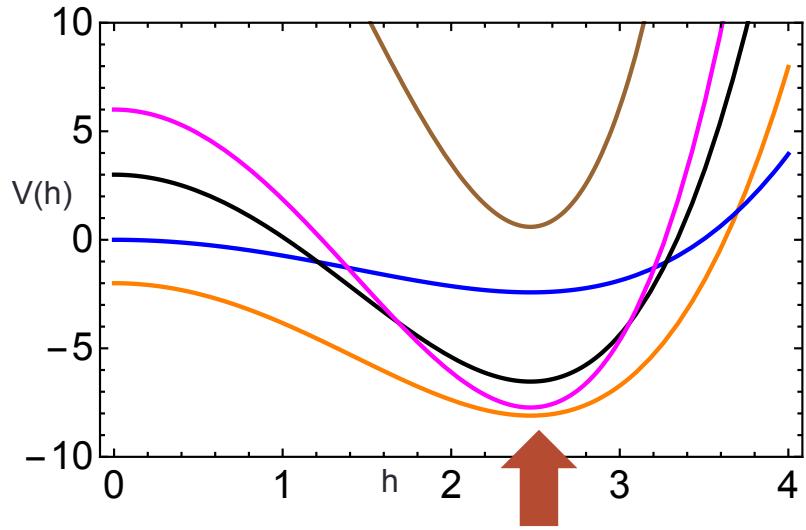
- Lowest energy state has $\langle h \rangle = v$
- This Higgs field value surrounds us all

What do we know about this potential?

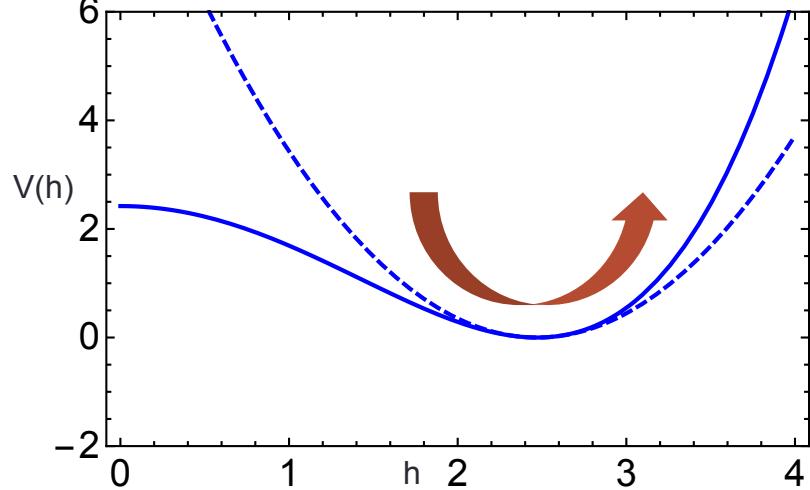
Classical potential: $V(h) = \Lambda + m^2 h^2 + \lambda h^4$

- 3 free parameters (Λ , m , λ)
 - Must be measured from data

Higgs potential

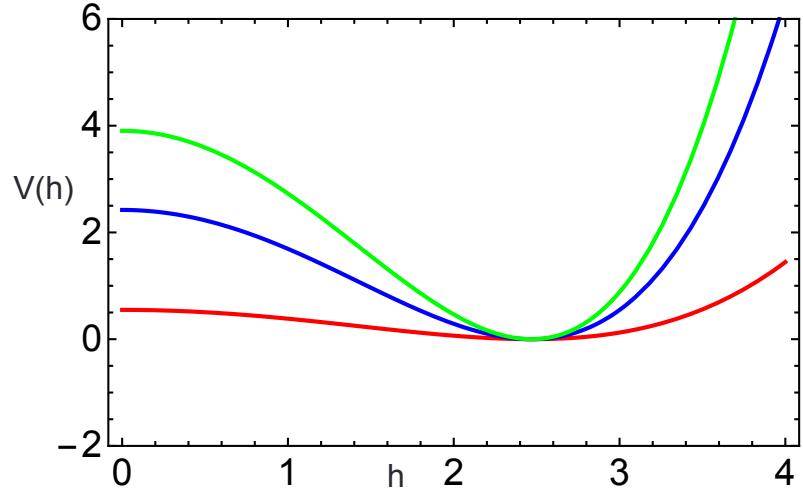


1933: Rate for beta decay ($G_F = \langle v \rangle^{-2}$)
gives vacuum expectation value $v = \frac{m}{\sqrt{\lambda}}$

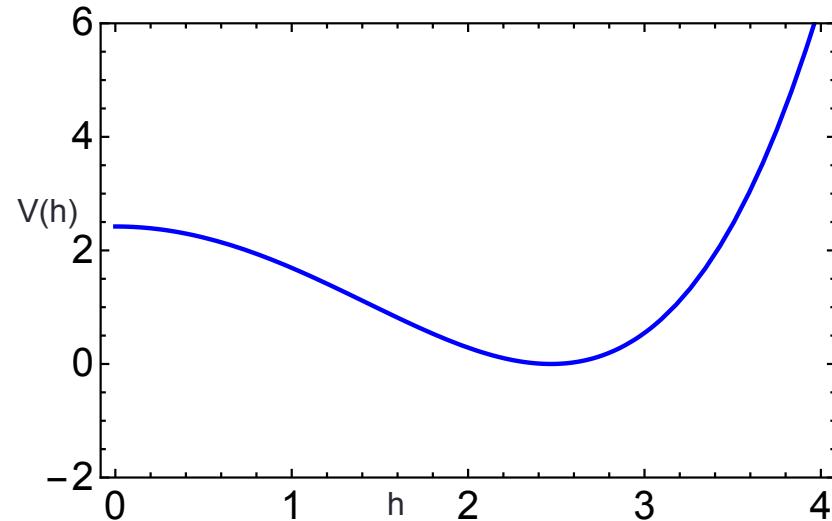


2012: Higgs boson mass $V''(v) = (126 \text{ GeV})^2$
gives curvature at minimum

$$V(h) = \Lambda + m^2 h^2 + \lambda h^4$$

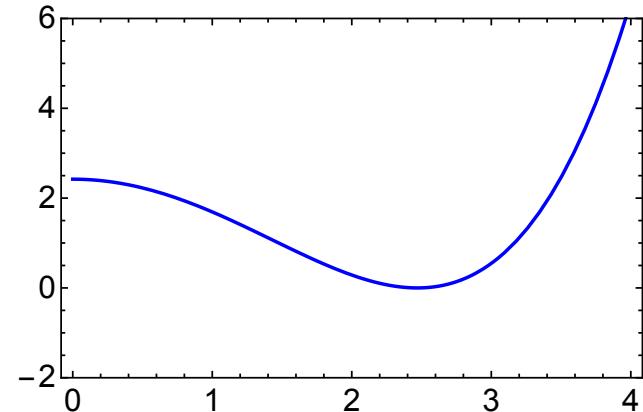


1998: acceleration of universe gives
vacuum energy density $V(v) = (10^{-3} \text{ eV})^4$



Classical potential: $V(h) = \Lambda + m^2 h^2 + \lambda h^4$

- 3 free parameters (Λ , m , λ)
 - Must be measured from data ✓

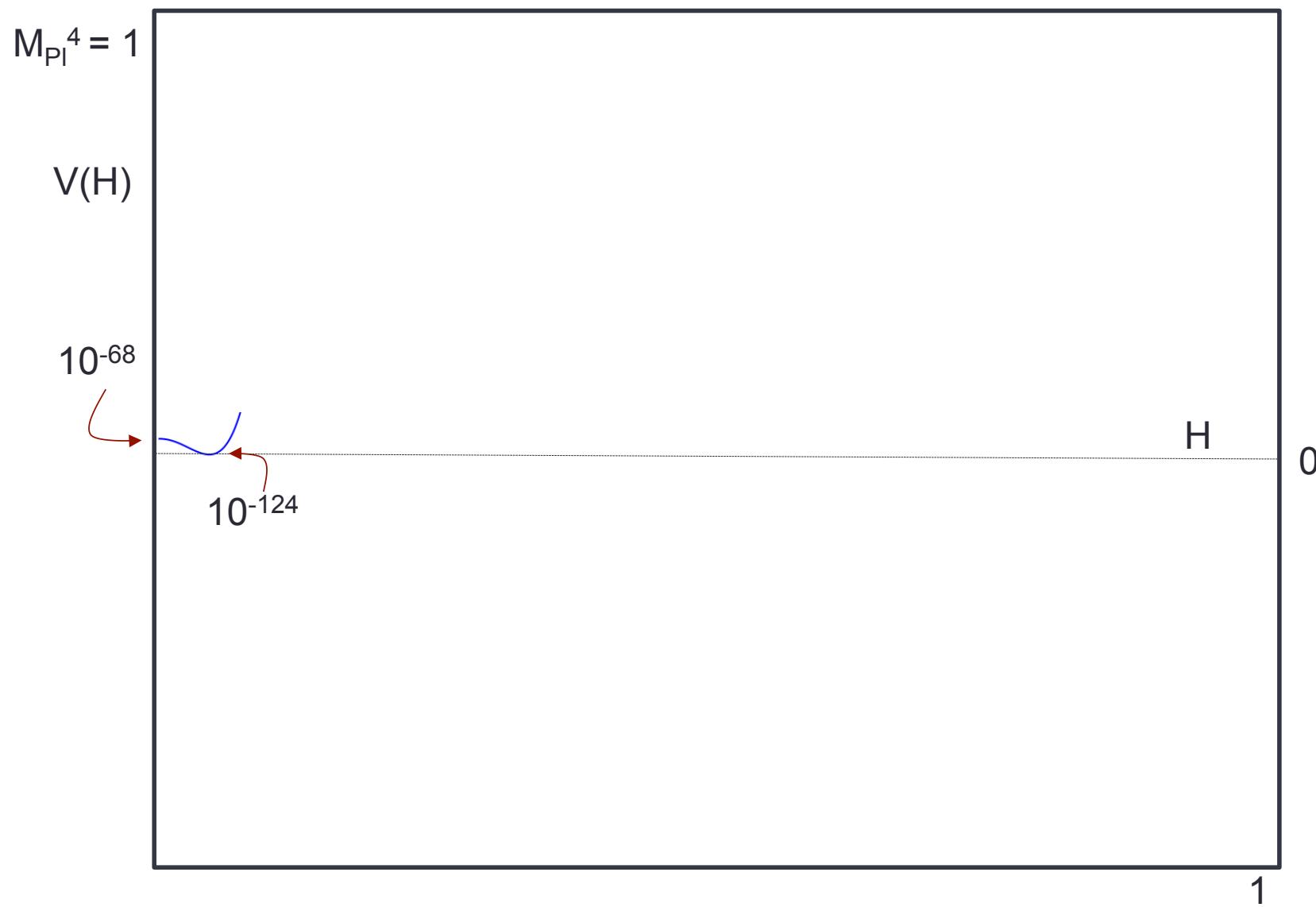


Why are the values of Λ , m , λ in nature **interesting**?

1. Fine tuning

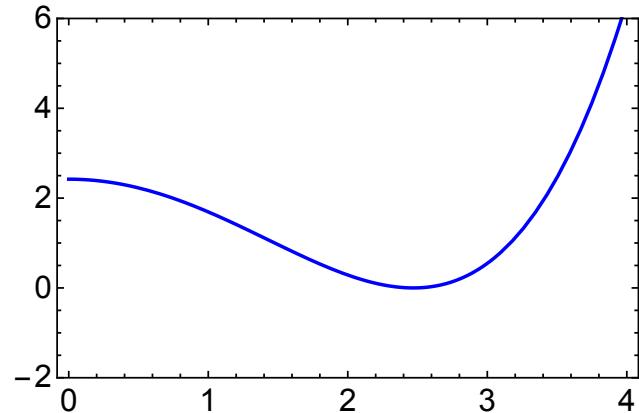
2. Vacuum stability

1. Fine tuning



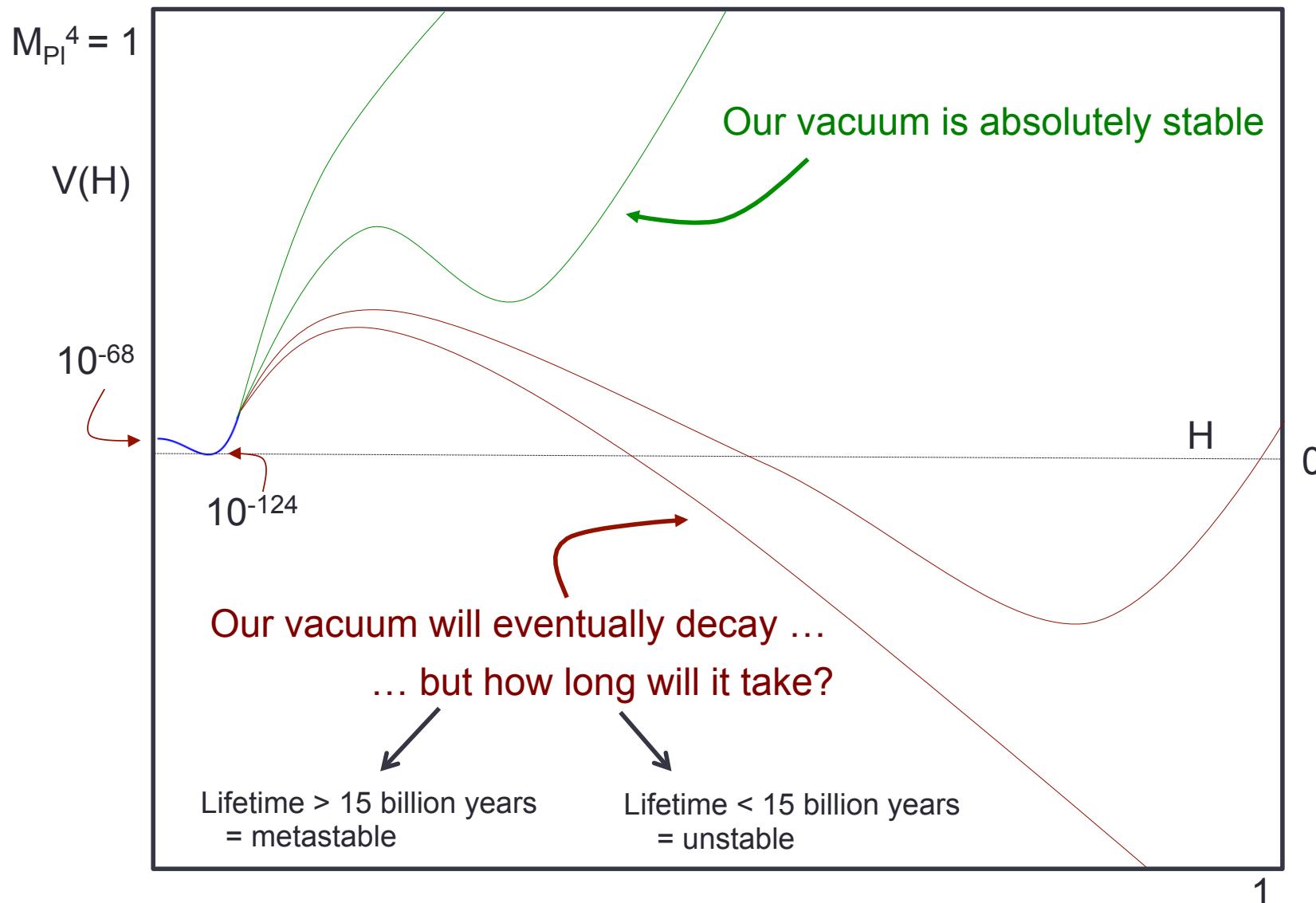
Classical potential: $V(h) = \Lambda + m^2 h^2 + \lambda h^4$

- 3 free parameters (Λ , m , λ)
 - Must be measured from data ✓
- **Only** 3 free parameters
 - Quantum Field Theory determines $V(h)$ for arbitrarily large h
 - Called the quantum-corrected or **Effective Potential**

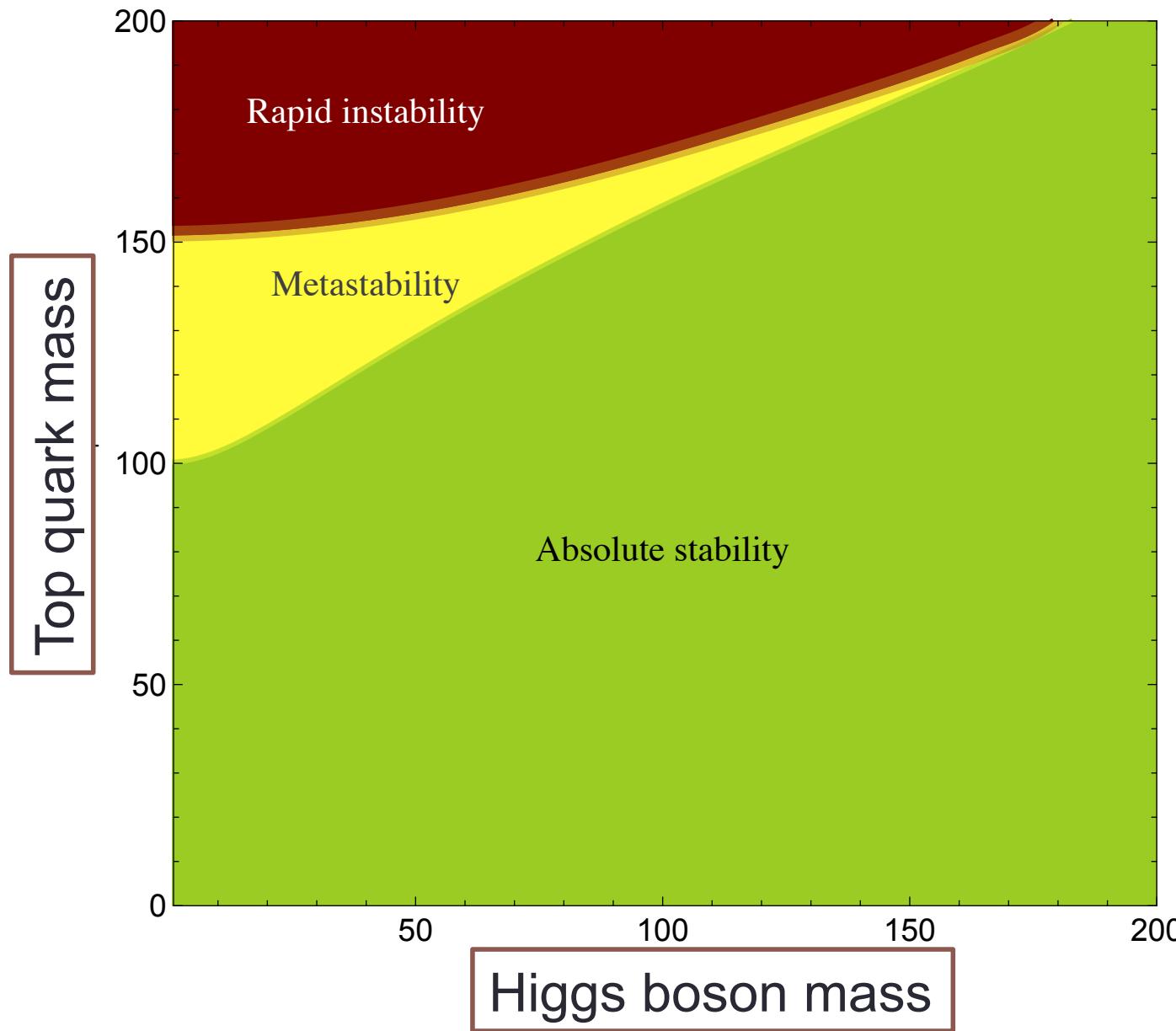


Fine tuning

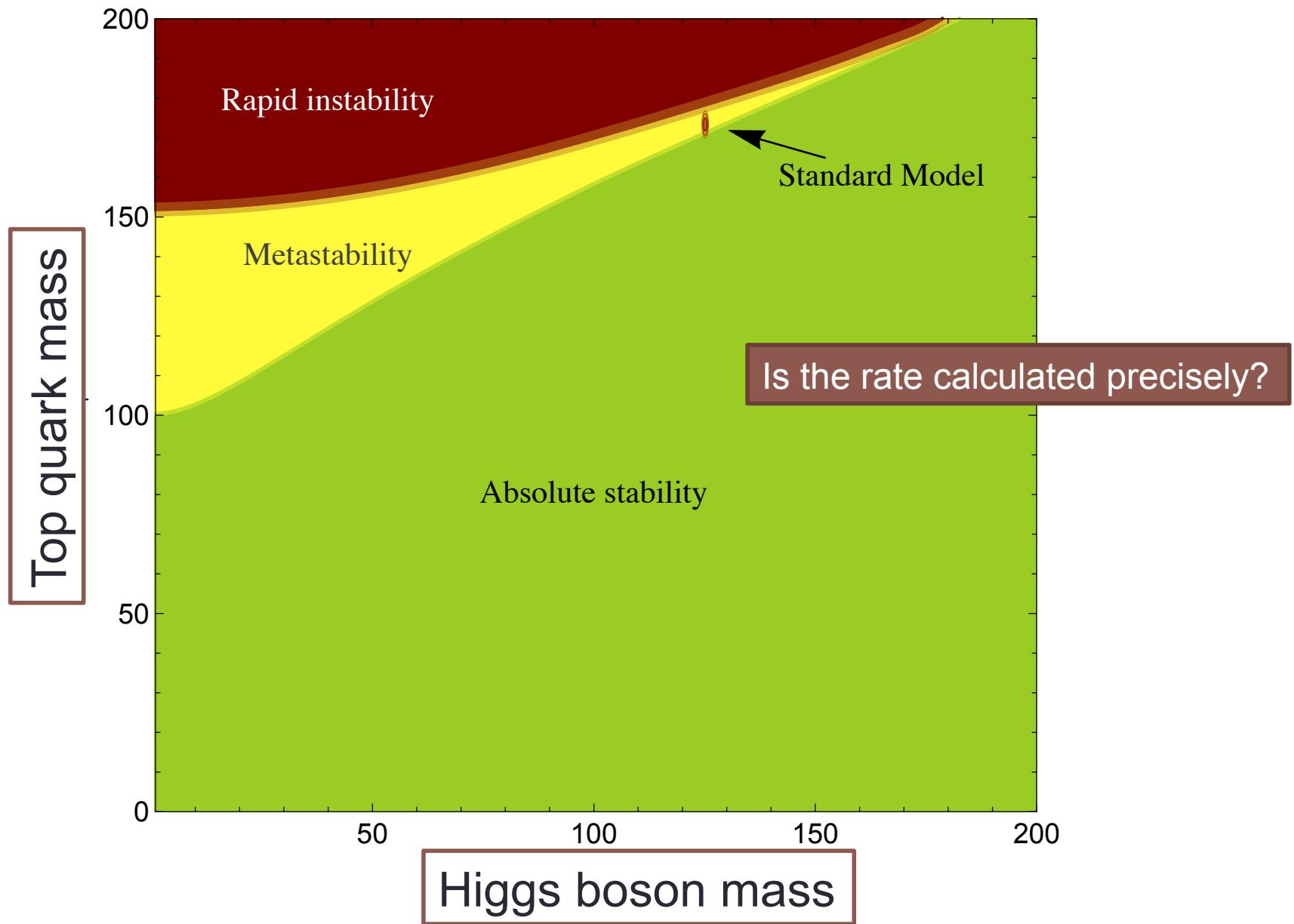
Stability



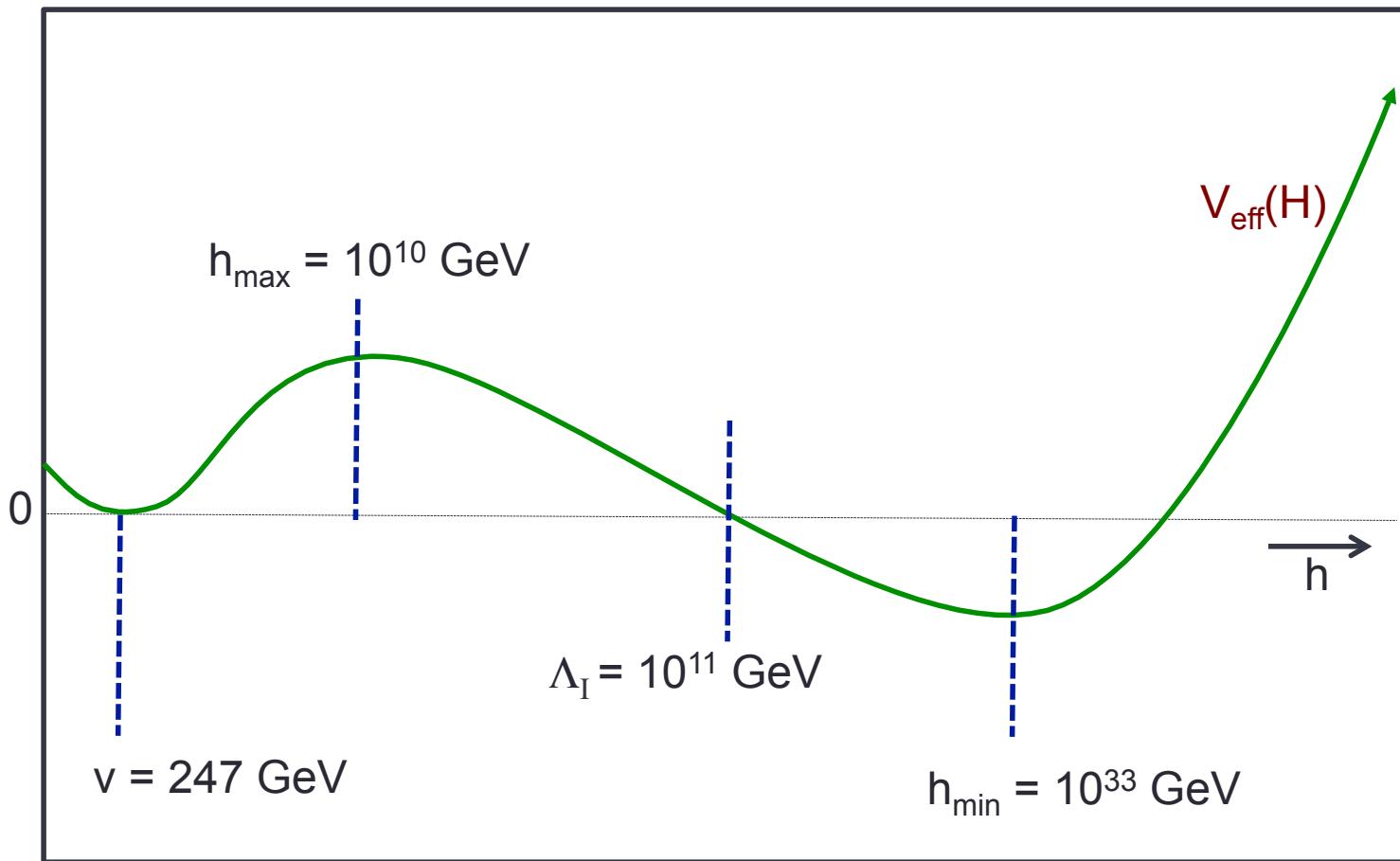
Stability phase diagram



Stability phase diagram



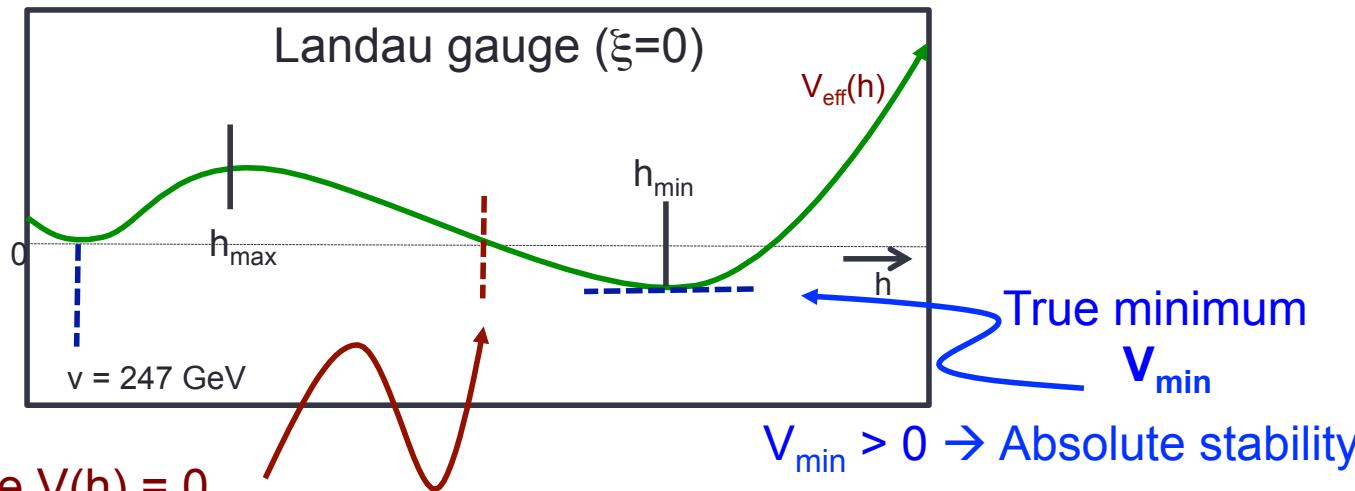
Standard Model Effective Potential



Are these scales physical?

Is the stability Planck sensitive?

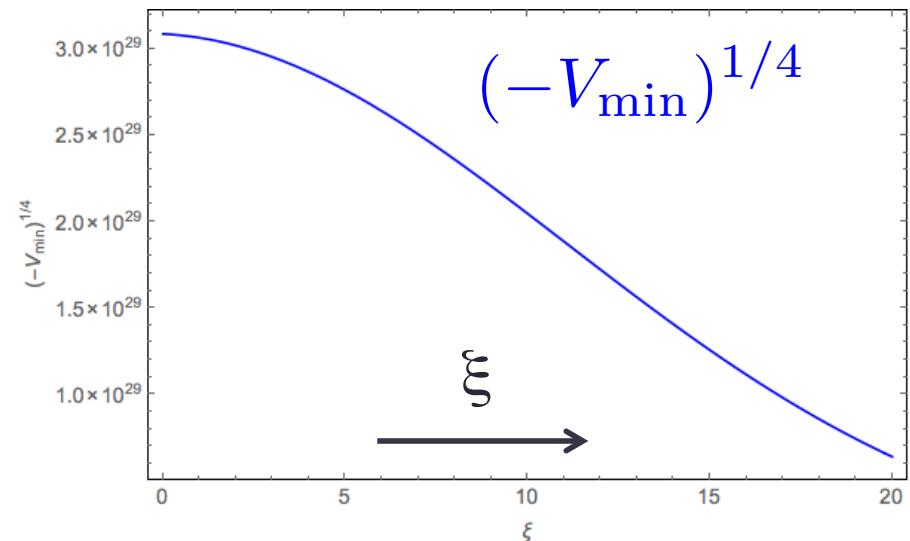
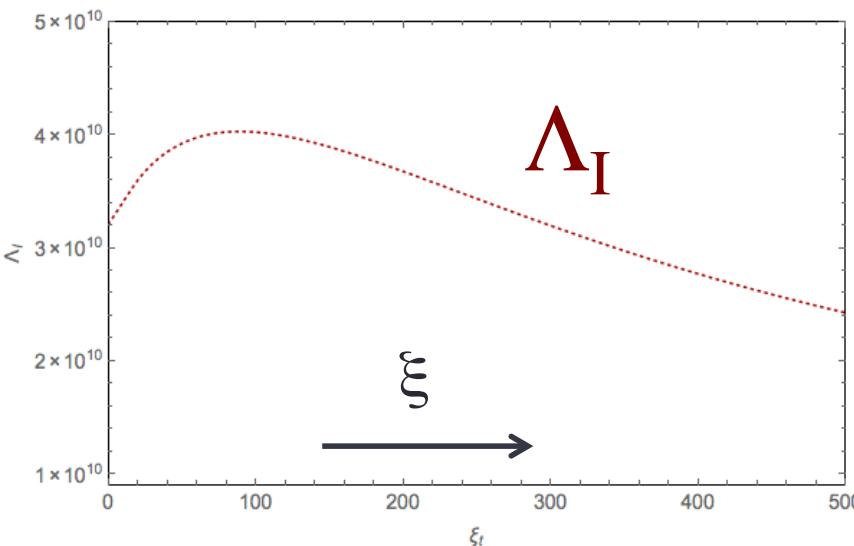
Gauge dependence



Instability scale Λ_I

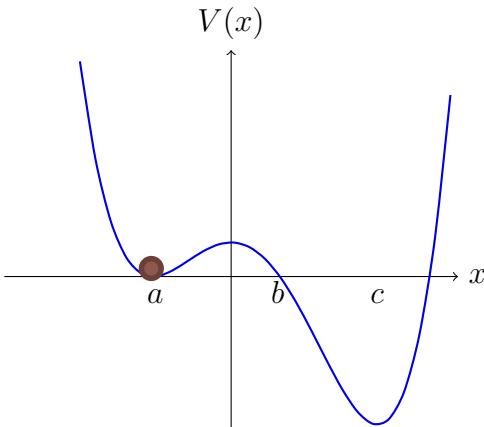
= value of h where $V(h) = 0$

- Indicates sensitivity to new physics



- h_{min} also gauge dependent
- h_{max} also gauge dependent
-

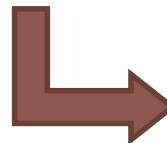
How do we calculate a decay rate?



Isolate ground state energy
from late times

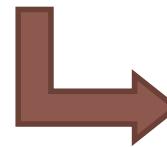
$$Z \equiv \langle a | e^{-H\mathcal{T}} | a \rangle = \int_{x(0)=a}^{x(\mathcal{T})=a} \mathcal{D}x e^{-S_E[x]}$$

$$Z = \sum_E e^{-ET} |\psi_E(a)|^2$$



$$E_0 = - \lim_{\mathcal{T} \rightarrow \infty} \frac{1}{\mathcal{T}} \ln Z$$

Decay rate is the imaginary
part of the energy



$$\frac{\Gamma}{2} = \text{Im} \lim_{\mathcal{T} \rightarrow \infty} \frac{1}{\mathcal{T}} \ln Z$$

Clearly this is not exactly what is meant

- **Z is real**
- True ground state at $E_c = V(c)$ has nothing to do with the false vacuum

How do we get an imaginary part?

Outline

Decay rates in QM and QFT
(1602.01102/1604.06090)

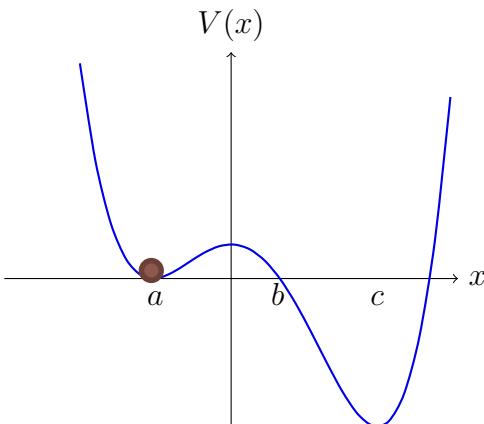
Effective potentials
(1408.0287 & 1408.0292)

1. The Coleman-Callan **potential-deformation method**
$$E = E_0 - \frac{1}{2} i \Gamma$$
 - What is E the energy of?
 - How does something manifestly real becomes complex?
2. **Solve the Schrodinger equation**
 - What exactly do we mean by a tunneling rate?
 - The two relevant time scales
 - The Gamow-Siegert prescription
3. **A direct approach**
 - Calculate the probability of going through the barrier
4. Using effective potentials to calculate decay rates
 - Resolving the gauge dependence issue
 - Understanding Planck sensitivity

1. THE POTENTIAL DEFORMATION METHOD

Coleman & Callan (1977)

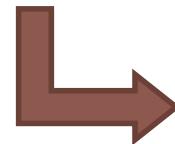
Coleman and Callan



$$Z \equiv \langle a | e^{-H\mathcal{T}} | a \rangle = \int_{x(0)=a}^{x(\mathcal{T})=a} \mathcal{D}x e^{-S_E[x]}$$

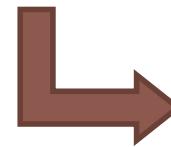
$$Z = \sum_E e^{-ET} |\psi_E(a)|^2$$

Isolate ground state energy
from late times



$$E_0 = - \lim_{\mathcal{T} \rightarrow \infty} \frac{1}{\mathcal{T}} \ln Z$$

Decay rate is the imaginary
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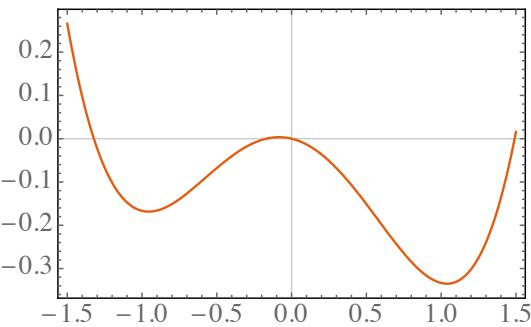
$$\frac{\Gamma}{2} = \text{Im} \lim_{\mathcal{T} \rightarrow \infty} \frac{1}{\mathcal{T}} \ln Z$$

Clearly this is not exactly what is meant

- **Z is real**
- True ground state at $E_c = V(c)$ has nothing to do with the false vacuum

How do we get an imaginary part?

Saddle points



$$Z \equiv \langle a | e^{-H\mathcal{T}} | a \rangle = \int_{x(0)=a}^{x(\mathcal{T})=a} \mathcal{D}x e^{-S_E[x]}$$

Dominated by **saddle points**
 = solutions to the Euclidean equations of motion

$$S = \int dt \left[\frac{1}{2} (\partial_t x) - V(x) \right]$$

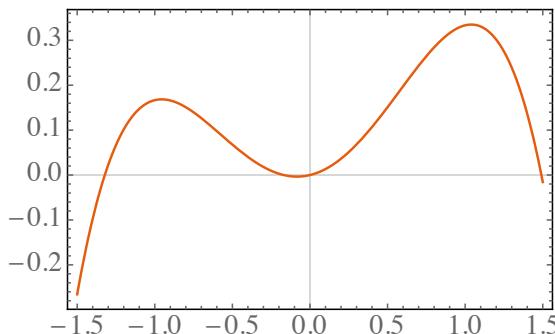
$$\partial_t^2 x = -V'(x)$$

$$S_E = \int d\tau \left[\frac{1}{2} (\partial_\tau x) + V(x) \right]$$

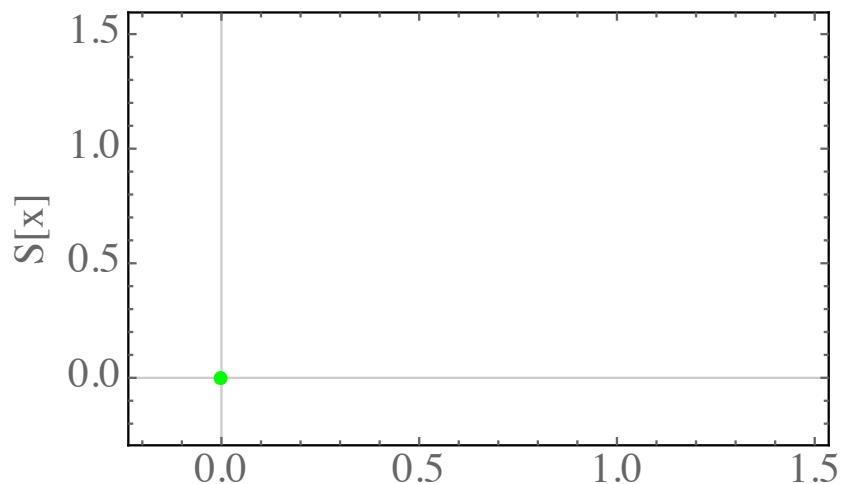
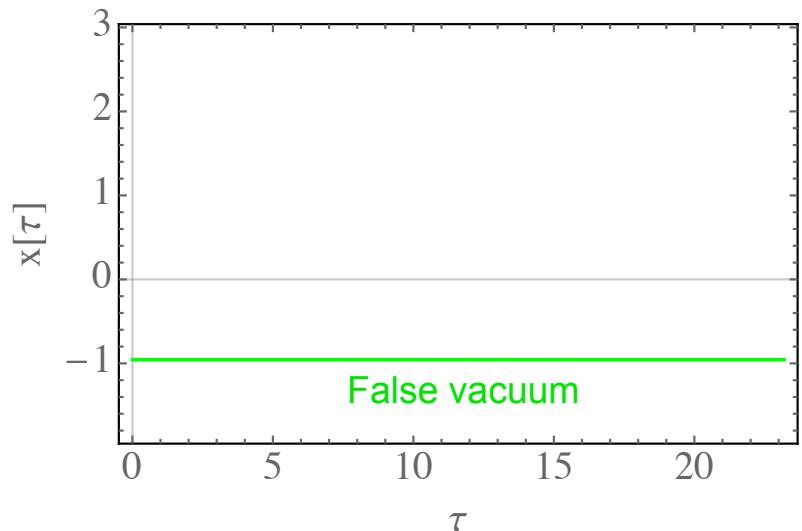
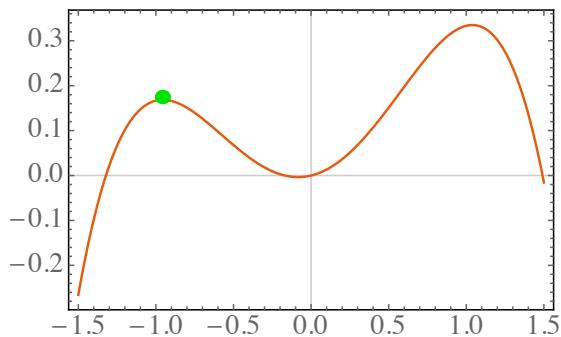
$$\partial_\tau^2 x = V'(x)$$



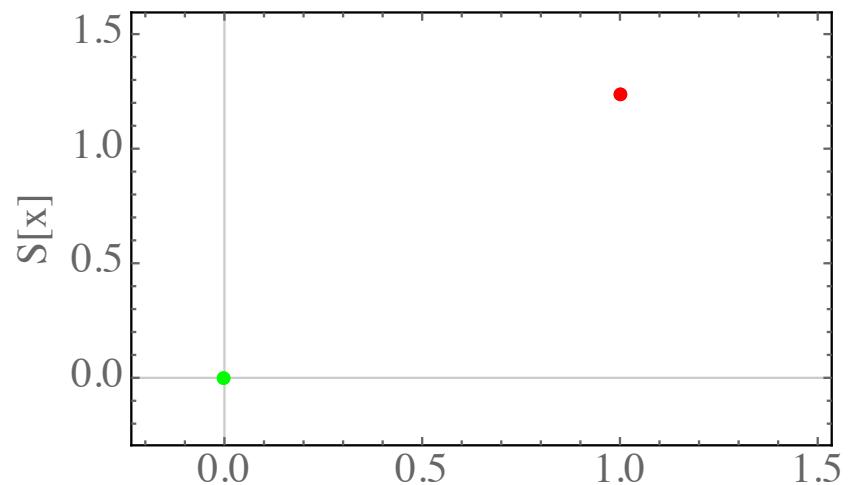
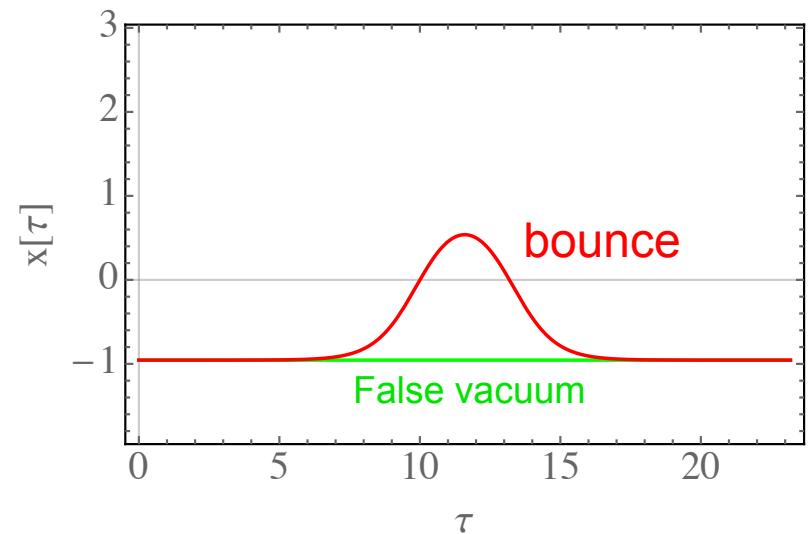
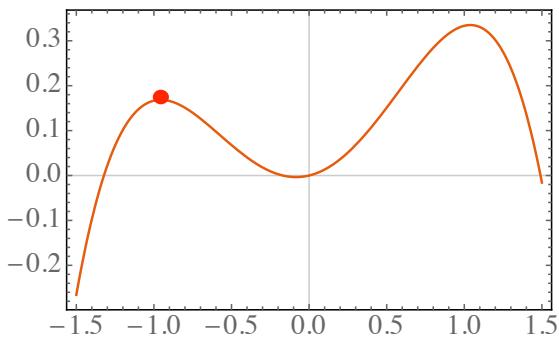
particle rolling down the **inverted potential**
 with boundary conditions $x(0) = x(\mathcal{T}) = a$



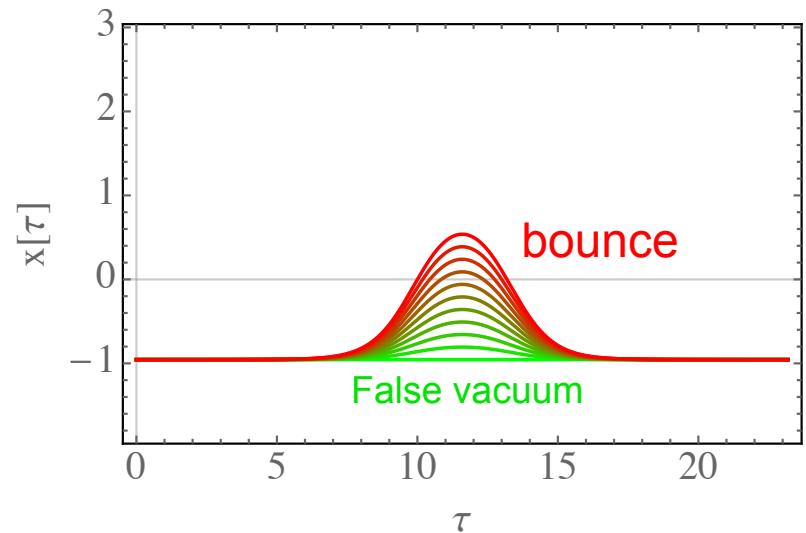
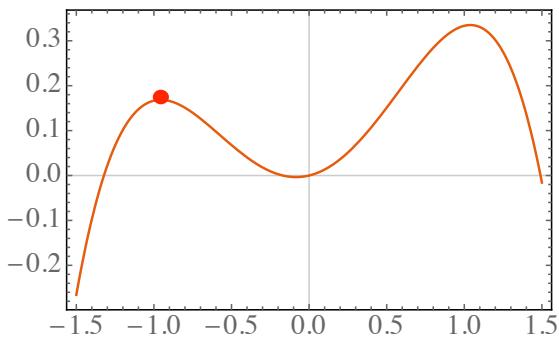
Saddle points



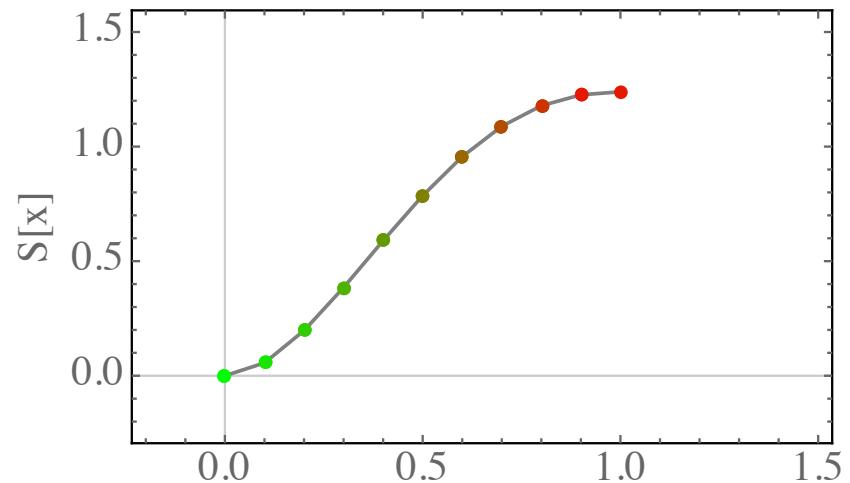
Saddle points



Saddle points



Bounce is a saddle point of action:
local maximum along one direction



Maximum \rightarrow negative eigenvalue of S'' \rightarrow Z has an imaginary part

$$\begin{aligned}
 Z &= \int_{x(0)=a}^{x(\mathcal{T})=a} \mathcal{D}x e^{-S_E[x]} \\
 &= \int_{x(0)=0}^{x(\mathcal{T})=0} \mathcal{D}x e^{-S_E[\bar{x}] - \frac{1}{2} S''_E[\bar{x}] x^2 - \dots} \\
 &= e^{-S_E[\bar{x}]} \int_{x(0)=0}^{x(\mathcal{T})=0} \mathcal{D}x e^{-\frac{1}{2} \int d\tau \{-x \partial_\tau^2 x + x V''(\bar{x}) x\}} \\
 &= \int d\xi_0 \dots d\xi_n e^{-\sum_n \frac{1}{2} \lambda_n \xi_n^2} \\
 &= \sqrt{\frac{2\pi}{\lambda_1}} \sqrt{\frac{2\pi}{\lambda_2}} \dots
 \end{aligned}$$

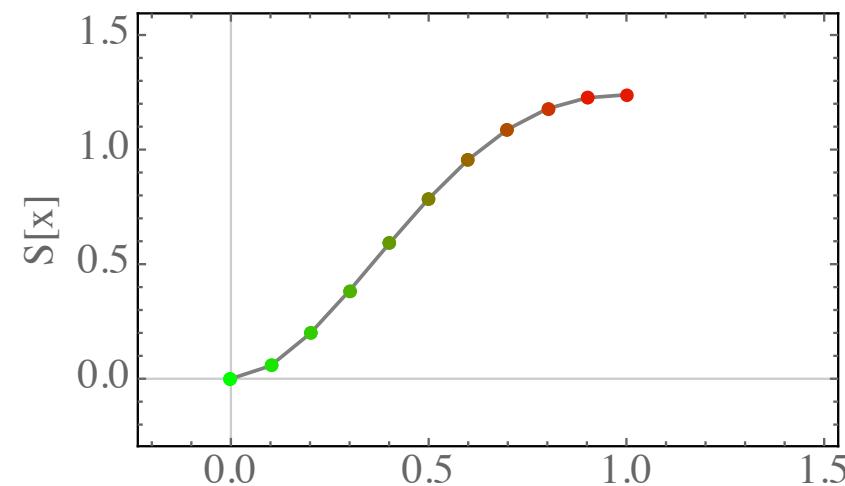
$$\frac{\Gamma}{2} = \text{Im} \lim_{\mathcal{T} \rightarrow \infty} \frac{1}{\mathcal{T}} \ln Z \neq 0$$

But Z is real! So how did this happen?

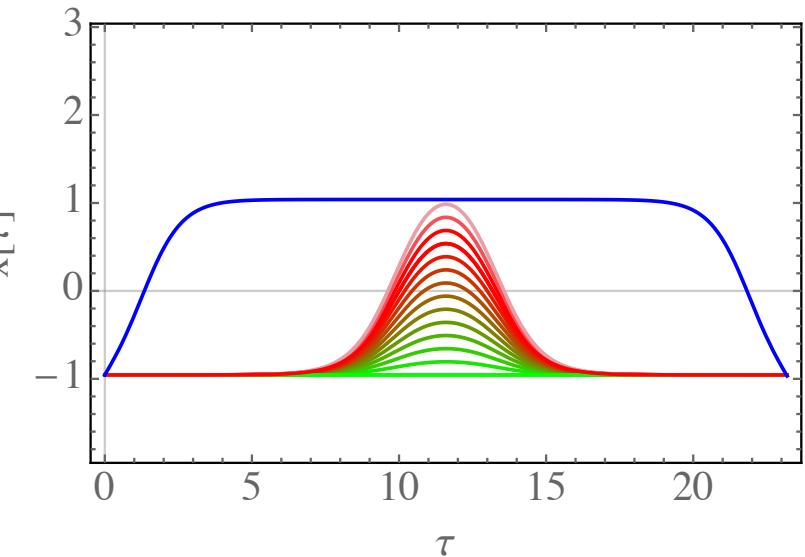
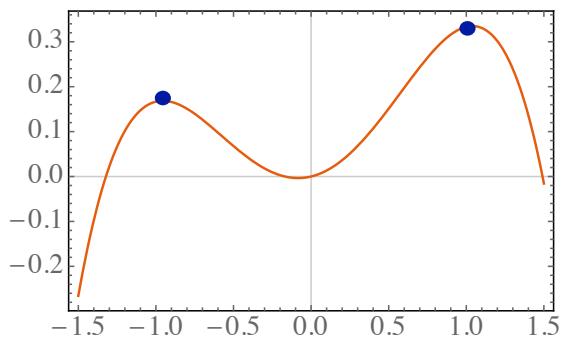
$$x(\tau) = \sum \xi_n y_n(\tau)$$

$$\begin{aligned}
 [-\partial_\tau^2 + V''(\bar{x})] y_n &= \lambda_n y_n \\
 \int d\tau y_n y_m &= \delta_{nm}
 \end{aligned}$$

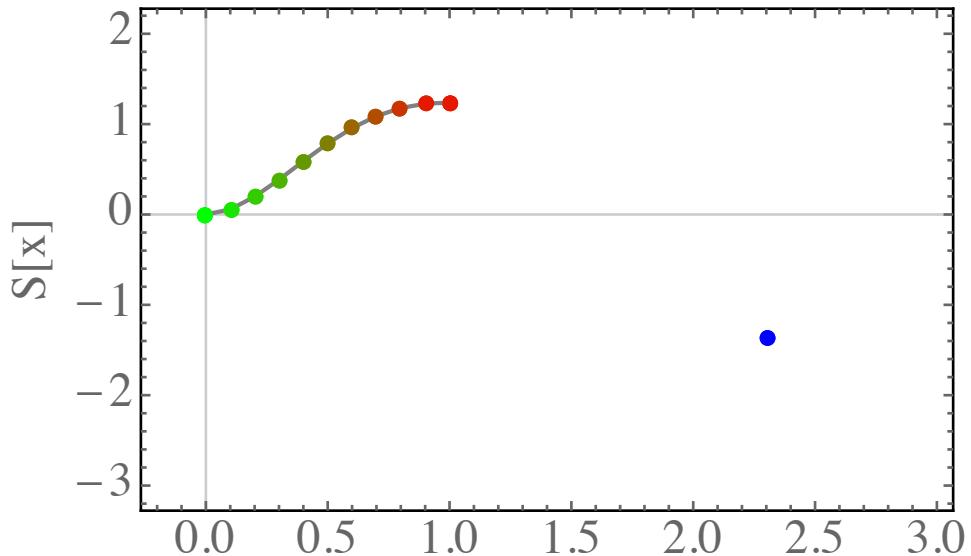
One of these is **negative ($\lambda_1 < 0$)**
if \bar{x} is a maximum of S in some direction



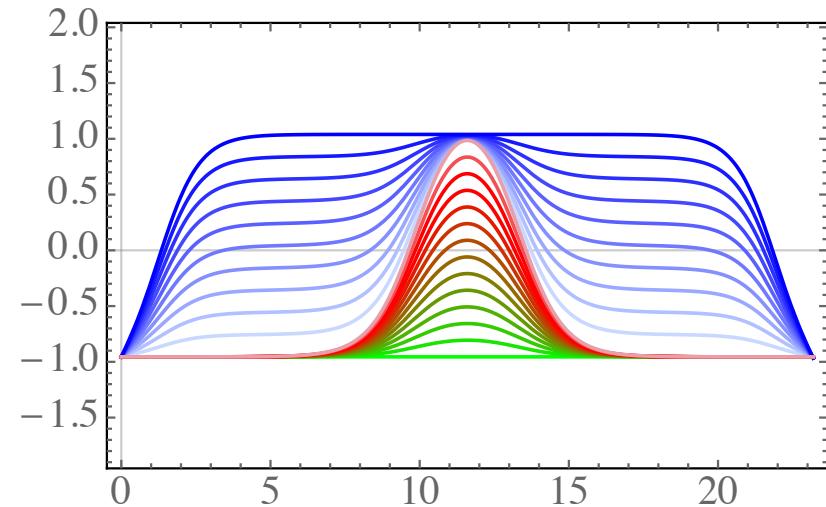
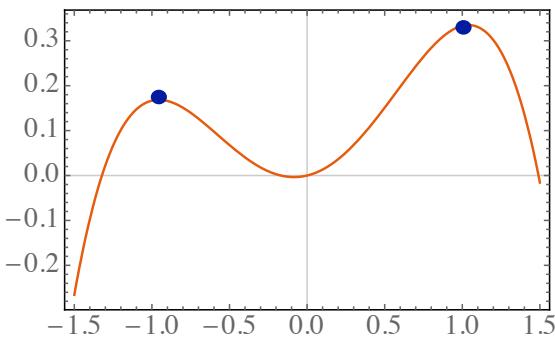
The Shot



Shot stays at true vacuum most of the time



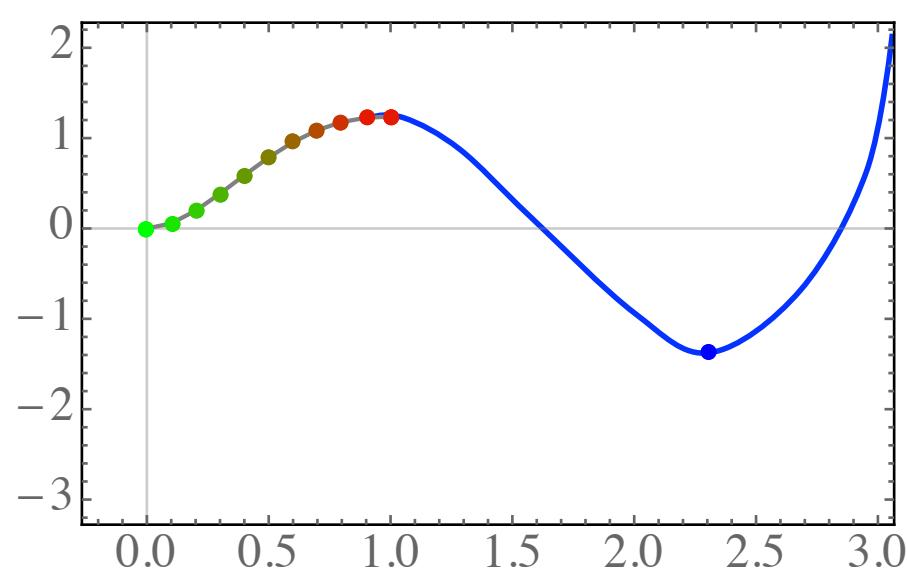
The Shot



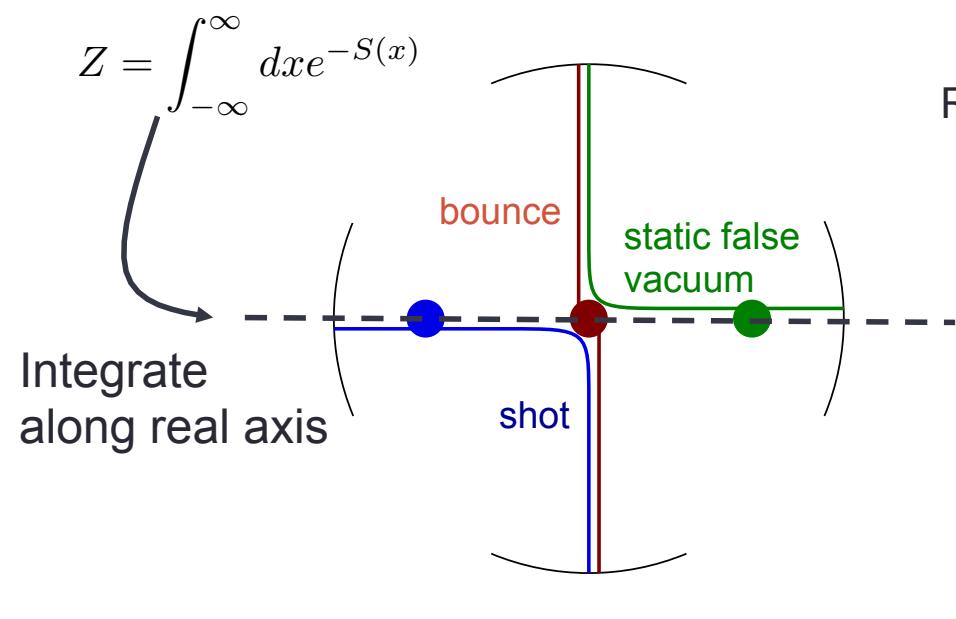
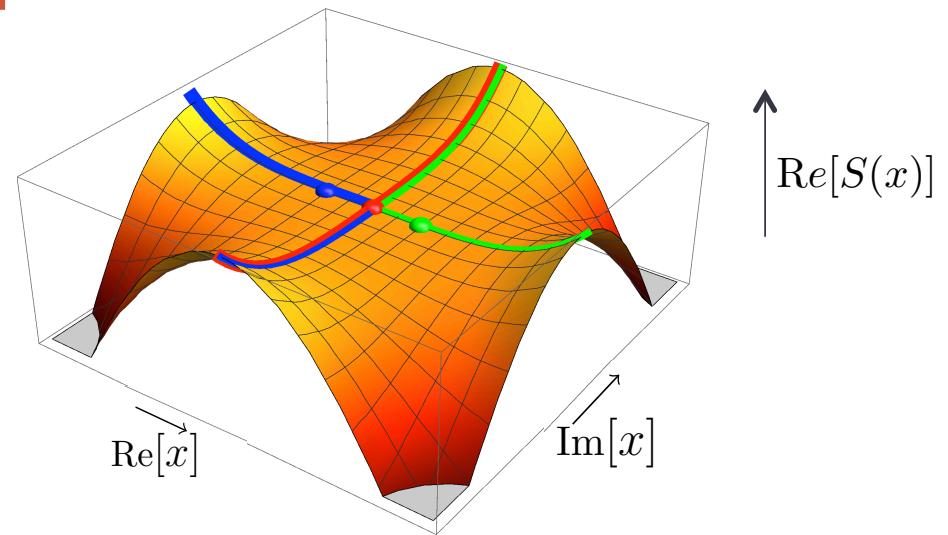
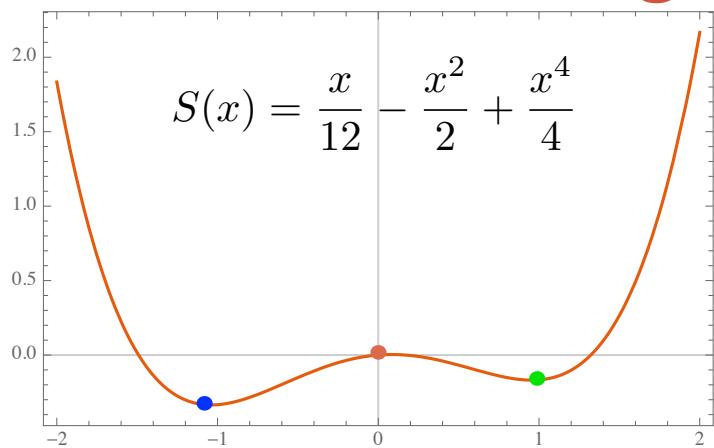
$$\begin{aligned}
 Z \equiv \langle a | e^{-H\mathcal{T}} | a \rangle &= \int_{x(0)=a}^{x(\mathcal{T})=a} \mathcal{D}x e^{-S_E[x]} \\
 &\approx e^{-S_E[x_{\text{shot}}]} \left(\gg e^{-S_E[x_{\text{bounce}}]} \right) \\
 &= e^{-E_0 \mathcal{T}}
 \end{aligned}$$

- **Bounce** is exponentially subdominant
- Consistent expansion must drop it
- **True vacuum dominates**

$S[x]$



Contour integration



Real axis = sum of steepest descent contours

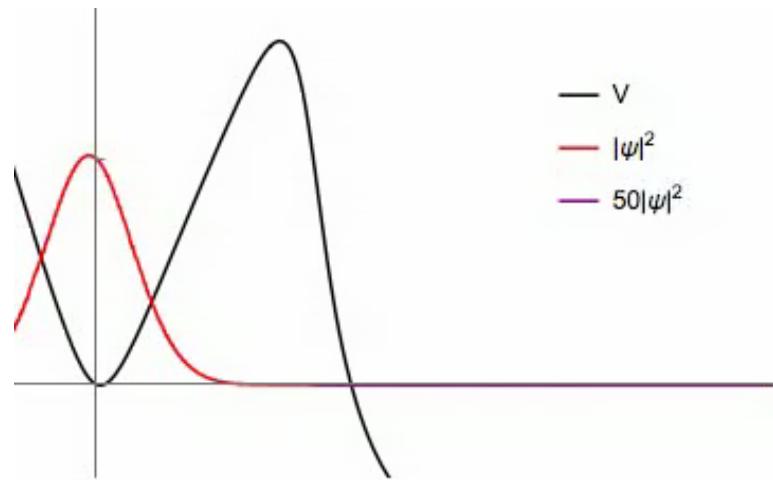
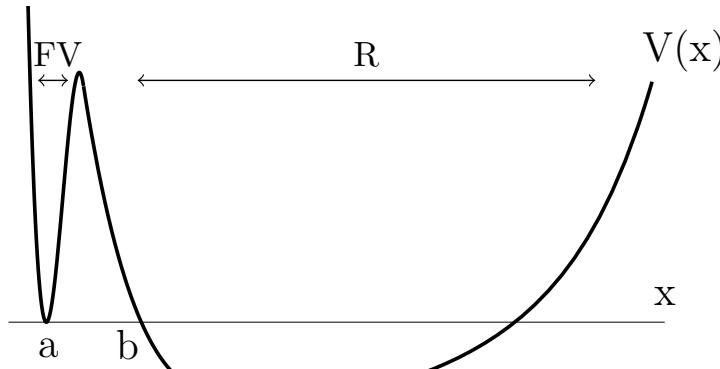
$$\begin{aligned}
 \int_{\mathbb{R}} &= \int_{C_{\text{shot}}} + \int_{C_{\text{bounce}}} + \int_{C_{\text{FV}}} \\
 &= e^{-S(x_{\text{shot}})} + i \frac{\Gamma}{2} \mathcal{T} \\
 &\quad + e^{-S(x_{\text{bounce}})} - i \Gamma \mathcal{T} \\
 &\quad + e^{-S(x_{\text{FV}})} + i \frac{\Gamma}{2} \mathcal{T} \\
 &\approx e^{-S(x_{\text{shot}})}
 \end{aligned}$$

2. SCHRODINGER EQUATION

Gamow (1928) & Siegert (1939)

Quantum mechanics

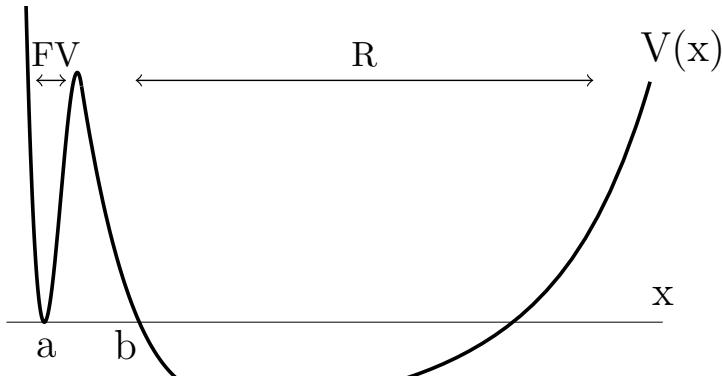
$$i\partial_t\psi(x, t) = \left[-\frac{1}{2m}\partial_x^2 + V(x)\right]\psi(x, t)$$



$$P_{FV}(T) \equiv \int_{FV} dx |\psi(x, T)|^2$$

Quantum mechanics

$$i\partial_t\psi(x, t) = \left[-\frac{1}{2m}\partial_x^2 + V(x)\right]\psi(x, t)$$

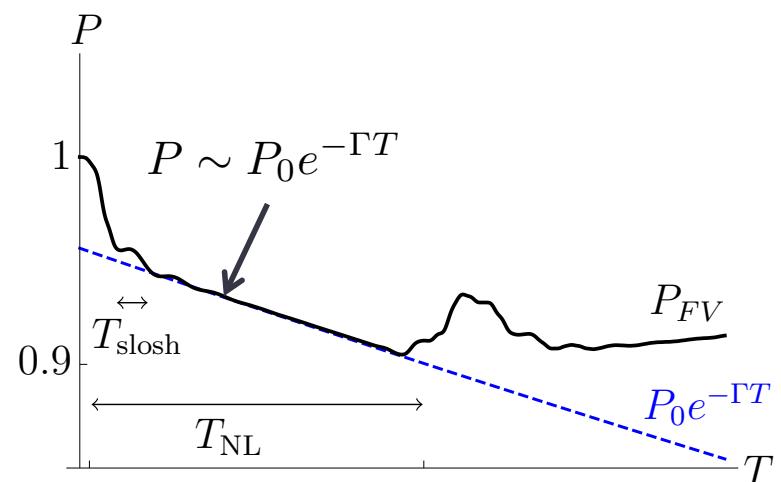
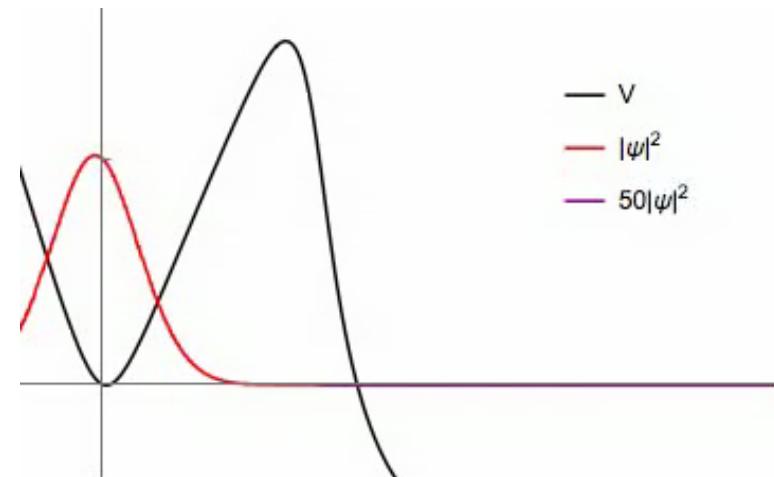


$$P_{\text{FV}}(T) \equiv \int_{\text{FV}} dx |\psi(x, T)|^2$$

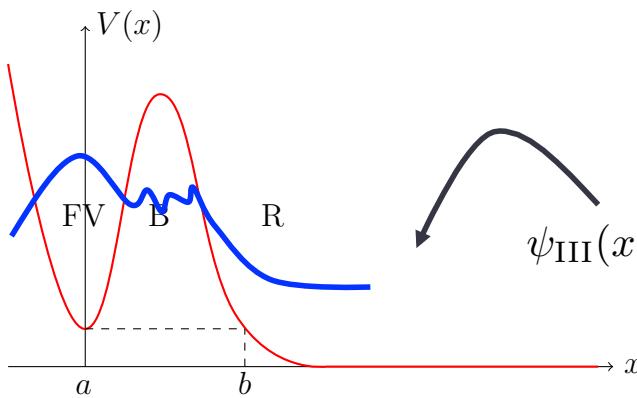
Two time scales

- $T > T_{\text{slosh}}$ – removes transients
- $T < T_{\text{NL}}$ -- avoids all ψ in true vacuum

$$\Gamma = - \lim_{\frac{T}{T_{\text{slosh}}} \rightarrow \infty} \lim_{\frac{T}{T_{\text{NL}}} \rightarrow 0} \frac{1}{P_{\text{FV}}} \frac{d}{dT} P_{\text{FV}}$$



Gamow's method



- Hermitian Hamiltonian \rightarrow energies are real
 $\rightarrow \psi^* \psi$ independent of time

$$\psi_{\text{III}}(x, t) = C e^{i(kx - Et)} + D e^{-i(kx - Et)}$$

Enforces $T \ll T_{\text{NL}}$ (no return flux)

Choose outgoing boundary conditions: $D=0$, $\psi_{\text{III}}(x, t) = C e^{i(kx - Et)}$

- Modes now have outgoing flux

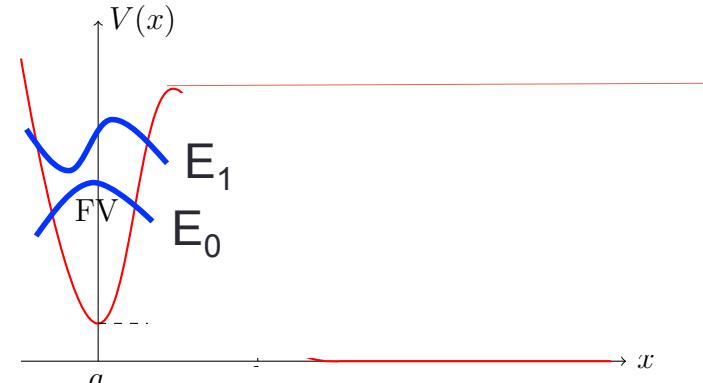
$$J = i(\psi^* \partial_x \psi - \psi \partial_x \psi^*) = -2p$$

- Violates unitarity \rightarrow energies are complex

$$E = E_0 - \frac{i}{2} \Gamma \quad \psi(x, t) = e^{-iE_0 t - \frac{1}{2} \Gamma t} \psi_0(x)$$

- Probability is time dependent

$$P = \int \psi^* \psi \sim e^{-\Gamma t}$$



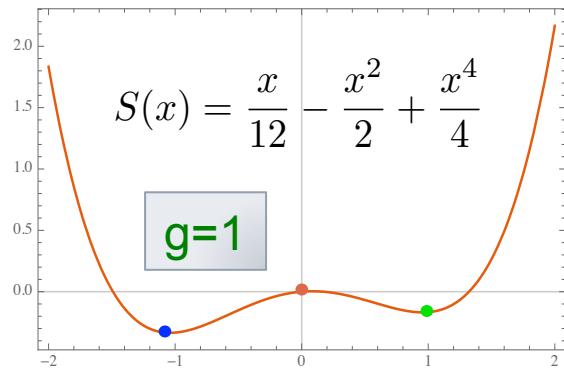
Assume E_1, E_2 etc components already died off

Enforces $T \gg T_{\text{slosh}}$ (only metastable FV decay)

1. THE POTENTIAL DEFORMATION METHOD (CONTINUED)

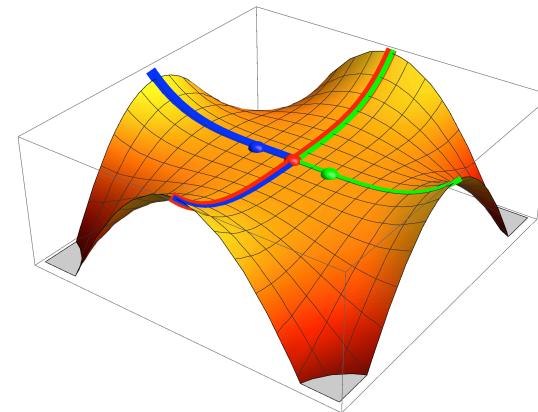
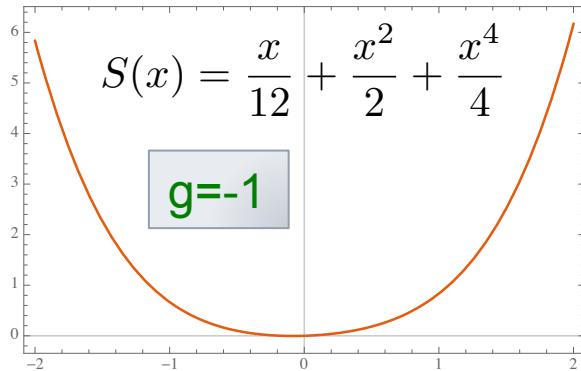
Coleman & Callan (1977)

Potential deformation

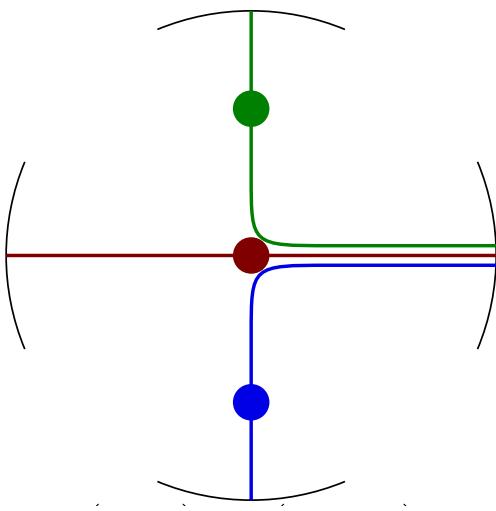
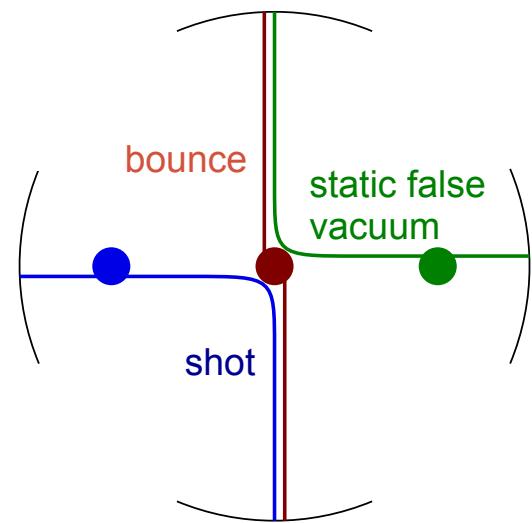
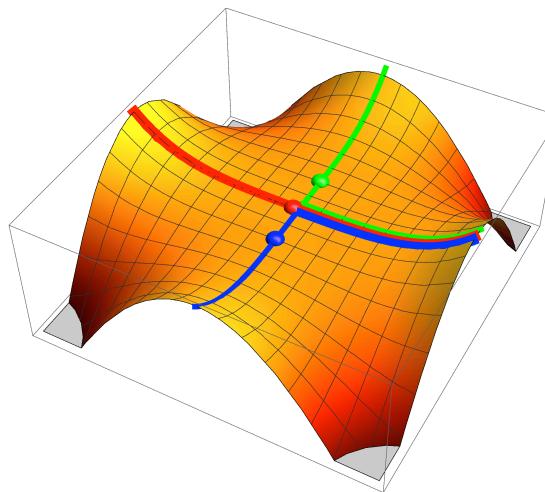


$$S_g(x) = \frac{x}{12} - g \frac{x^2}{2} + \frac{x^4}{4}$$

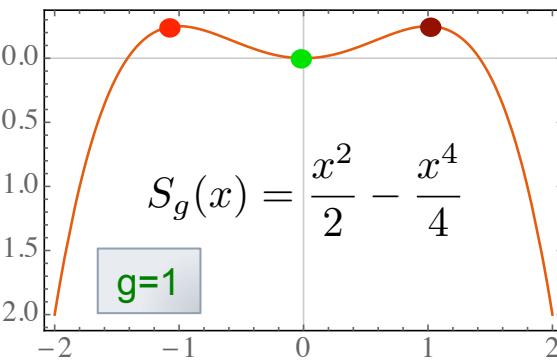
**deform potential
to prevent tunneling**



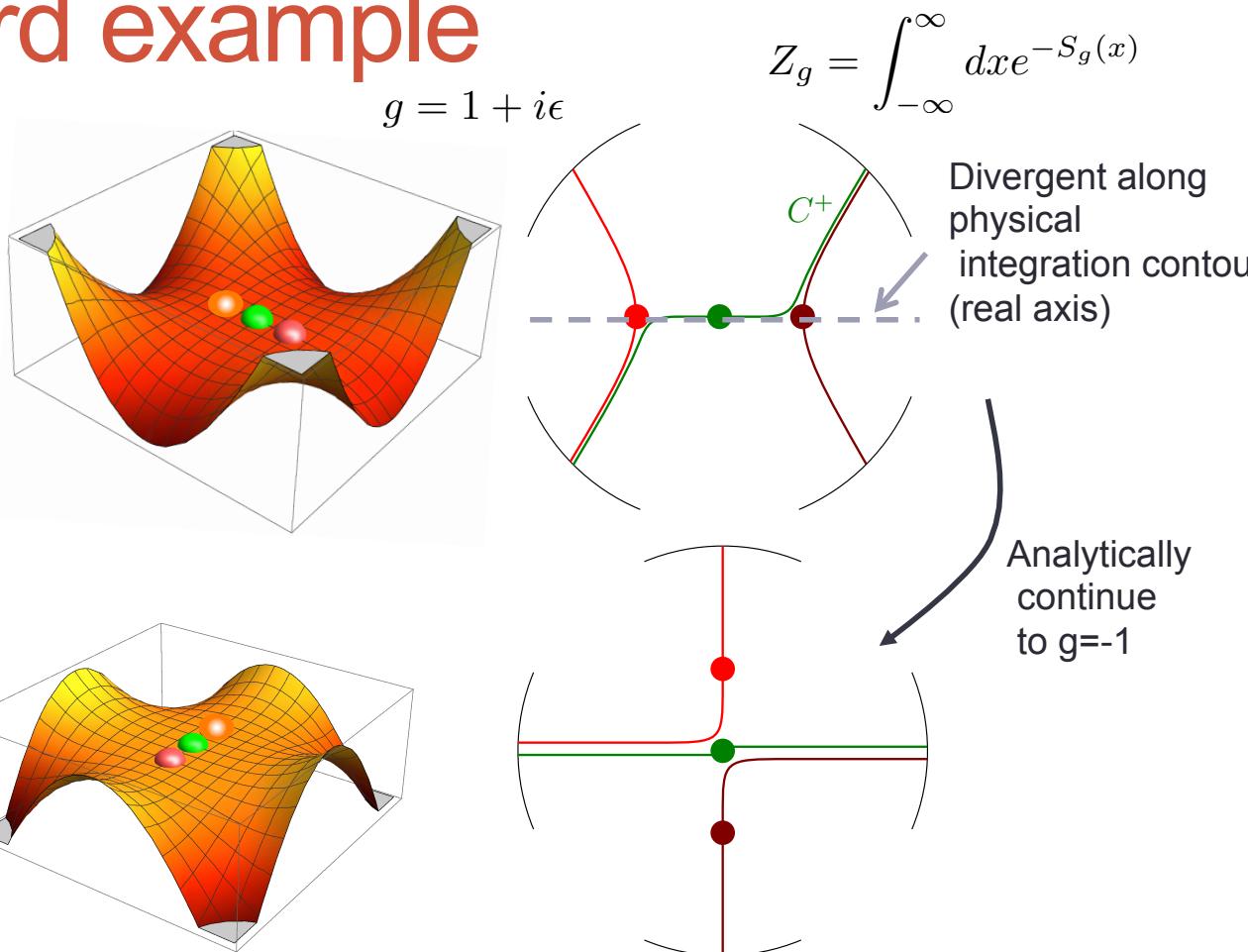
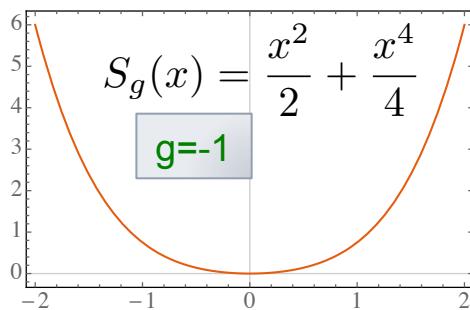
$$Z_g = \int_{-\infty}^{\infty} dx e^{-S_g(x)} \left\{ \begin{array}{l} \cdot \text{ real at } g = +1 \\ \cdot \text{ real at } g = -1 \\ \cdot \text{ an analytic function of } g \end{array} \right.$$



More standard example



$$S_g(x) = \frac{x^2}{2} - g \frac{x^4}{4}$$



- Fix integration to be **along** contour passing through saddle at $x=0$
- Return to $g=1$, keeping integration **along** green contour
- Z now has imaginary part at $g=1$

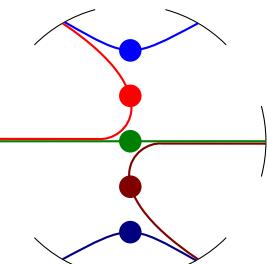
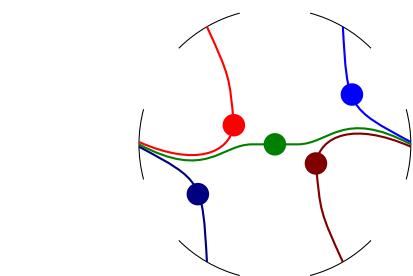
Well-defined procedure. But is the imaginary part the decay rate?

Add convergence factor

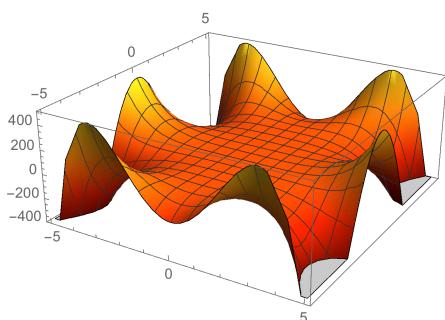
- Modifying potential/action away from region of interest should not affect rate

$$S_g(x) = \frac{x^2}{2} - g \frac{x^4}{4} + \frac{x^6}{60}$$

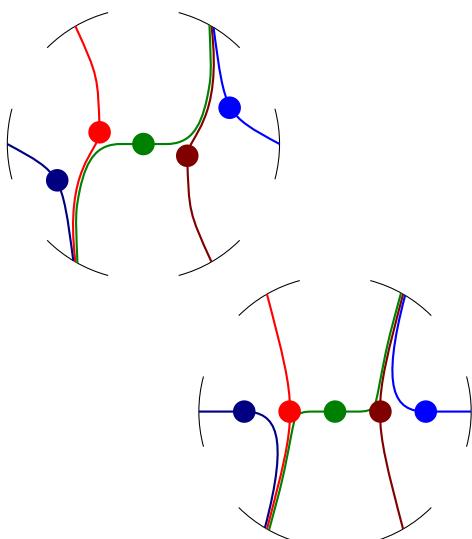
$$g = \exp(i\frac{\pi}{4})$$



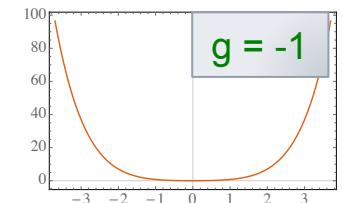
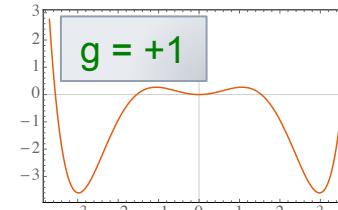
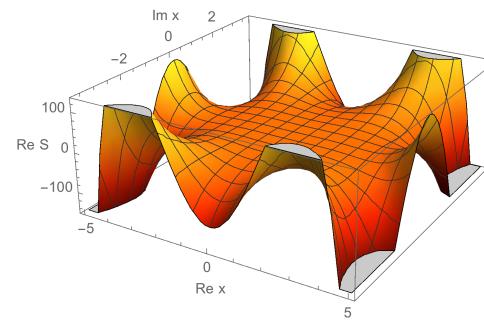
$$g = -\exp(-i\epsilon)$$



$$g = \exp(i\frac{\pi}{3})$$



$$g = \exp(i\epsilon)$$



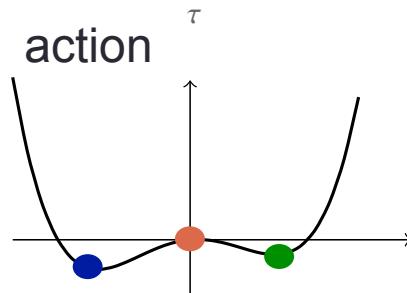
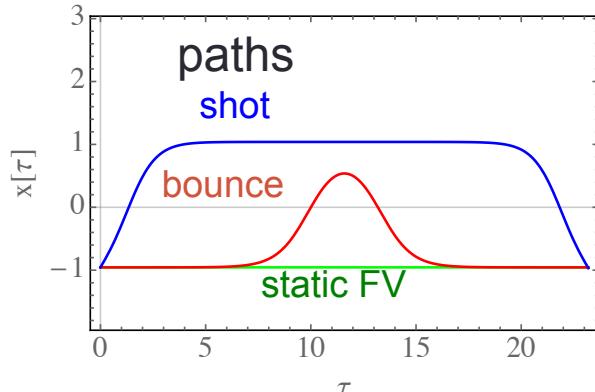
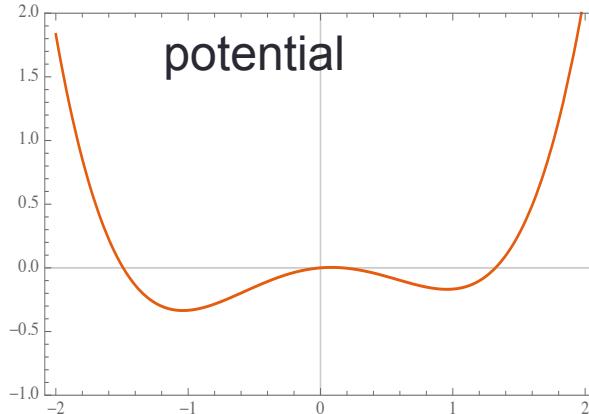
$$Z_g = \int_{-\infty}^{\infty} dx e^{-S_g(x)}$$

- real at $g = +1$
- real at $g = -1$
- Z_g is an analytic function of g

- We can still fix the contour at $g = -1$ and follow it back.

But why?

Physical limits



$T \gg T_{\text{slosh}}$ (only metastable FV decay)

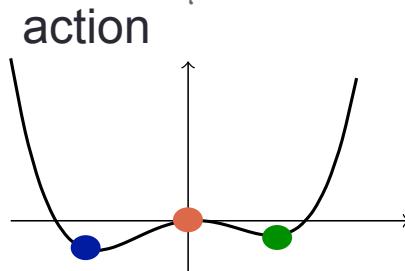
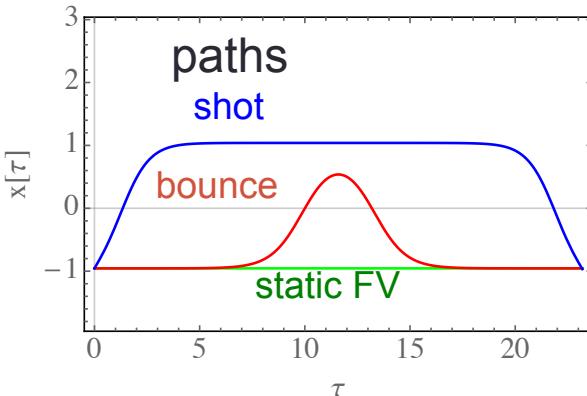
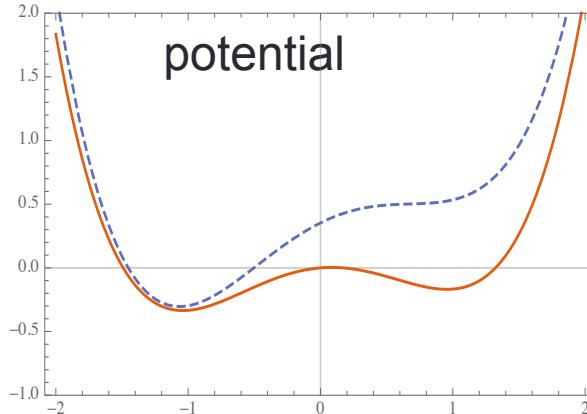
$T \ll T_{\text{NL}}$ (no return flux)

$$\Gamma = - \lim_{\frac{T}{T_{\text{slosh}}} \rightarrow \infty} \lim_{\frac{T}{T_{\text{NL}}} \rightarrow 0} \frac{1}{P_{\text{FV}}} \frac{d}{dT} P_{\text{FV}}$$

$$\begin{aligned} Z &\equiv \langle a | e^{-H\mathcal{T}} | a \rangle = \int_{x(0)=a}^{x(\mathcal{T})=a} \mathcal{D}x e^{-S_E[x]} \\ &\sim e^{-E_0 T} + e^{-E_{FV} T} \\ \xrightarrow{\text{L}} \quad & - \lim_{T \rightarrow \infty} \frac{1}{T} \ln Z = \min(E_0, E_{FV}) \end{aligned}$$

Taking $T \rightarrow \infty$ picks out true ground state E_0

Physical limits



$T \gg T_{\text{slosh}}$ (only metastable FV decay)

$T \ll T_{\text{NL}}$ (no return flux)

$$\Gamma = - \lim_{\frac{T}{T_{\text{slosh}}} \rightarrow \infty} \lim_{\frac{T}{T_{\text{NL}}} \rightarrow 0} \frac{1}{P_{\text{FV}}} \frac{d}{dT} P_{\text{FV}}$$

$$\begin{aligned} Z &\equiv \langle a | e^{-H\mathcal{T}} | a \rangle = \int_{x(0)=a}^{x(\mathcal{T})=a} \mathcal{D}x e^{-S_E[x]} \\ &\sim e^{-E_0 T} + e^{-E_{FV} T} \\ \Rightarrow & -\lim_{T \rightarrow \infty} \frac{1}{T} \ln Z = \min(E_0, E_{FV}) \end{aligned}$$

Taking $T \rightarrow \infty$ picks out true ground state E_0

We want to

1. Deform the potential so FV is true ground state
2. Take $T \rightarrow \infty$
 - Picks out $E_{FV}(g)$
3. Deform back

$T \ll T_{\text{NL}}$ (no return flux)

$T \gg T_{\text{slosh}}$ (only metastable FV decay)

The $T \rightarrow \infty$ limit **does not commute** with analytic continuation

- $\min(E_0, E_{FV})$ is **not analytic**

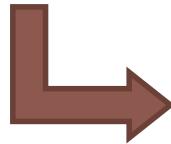
$$Z \sim \int dx e^{-S_E} \sim \int dx e^{-\frac{TE_{FV}}{\hbar}}$$

\leftarrow
 $T \rightarrow \infty$ limit
 like $\hbar \rightarrow 0$ limit
 forces saddle point approximation

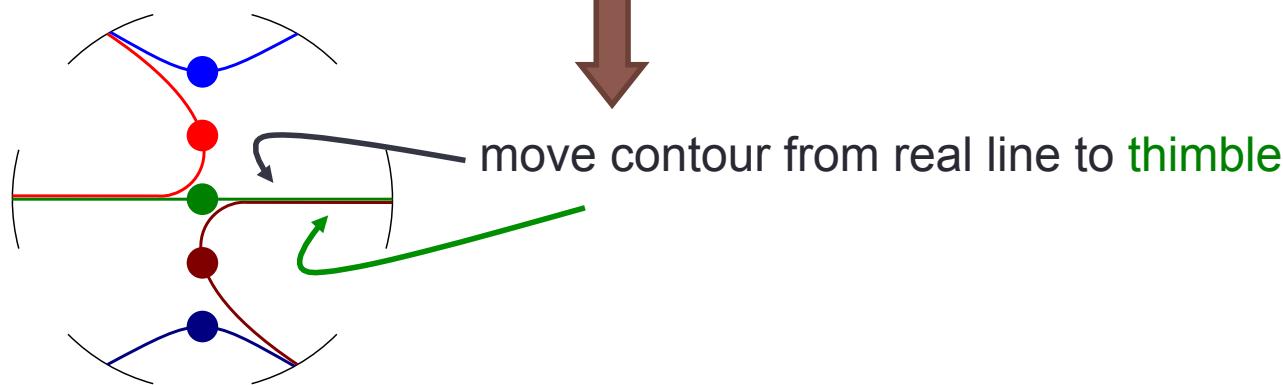
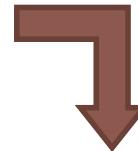
Saddle point approximation

$$\begin{aligned}
 Z_g &= \int dx e^{\frac{1}{\hbar}(-\frac{1}{2}x^2 + \frac{g}{4}x^4)} \\
 &\approx \int dx e^{-\frac{1}{2\hbar}x^2} \left[1 + \frac{g}{4}x^4 + \frac{1}{2} \left(\frac{g}{4}x^4 \right)^2 + \dots \right] \\
 &= \sqrt{2\pi} \left(1 + \frac{3g}{4} + \frac{105g^2}{32} + \frac{3465g^3}{128} + \frac{675\,675g^4}{2048} + \frac{43\,648\,605g^5}{8192} + \frac{7\,027\,425\,405g^6}{65\,536} + \dots \right)
 \end{aligned}$$

- Asymptotic series
- Coefficients grow factorially
- Summing the series does **not** reproduce the original function

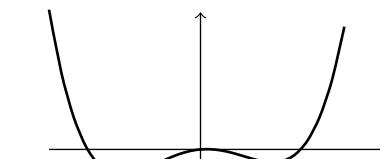


Performing the saddle point approximation does not commute with analytic continuation



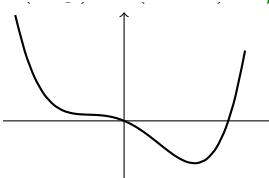
Examples

$$S_g(x) = h \frac{x}{12} - g \frac{x^2}{2} + \frac{x^4}{4}$$

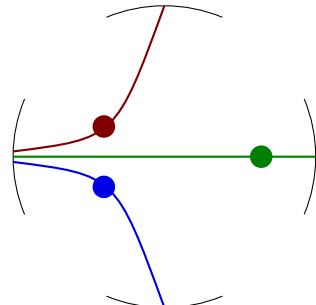


$$(h, g) = (1, 1 + i\epsilon)$$

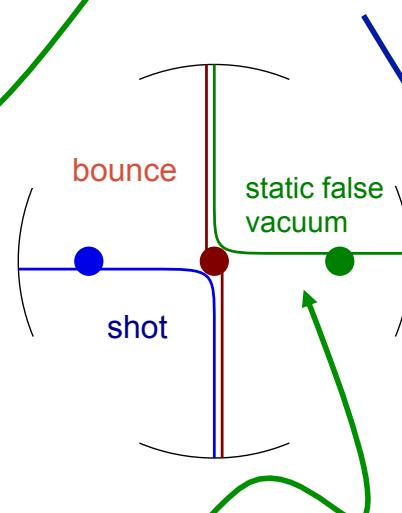
Deform to stabilize
false vacuum



$$(h, g) = (-5, 1)$$



$T \rightarrow \infty$ limit fixes to green contour

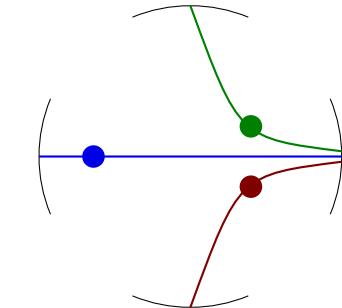
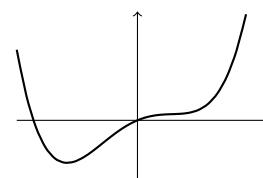


Z has imaginary part
equal to **half** of the
bounce contour

Deform to stabilize
bounce

Z has imaginary part
equal to **all** of the **bounce contour**

Deform to stabilize
shot



Z has imaginary part
equal to **minus half** of the **bounce contour**

- Probability grows with time

Discontinuity

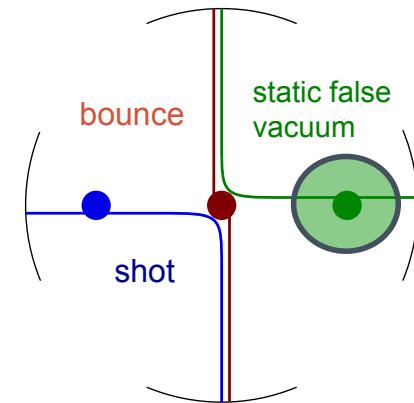
Can we just integrate along the **FV** contour?

Yes, at least for this toy integral

$$Z = \int_C dx e^{-S(x)}$$

No

- Not clear what “fixing to a contour” means for a path integral
- Saddle point approximation **loses the imaginary part**
 - Expanding around the saddle gives a **real integral**
 - Imaginary part comes from region far away
- Saddle point approximation **does work** for the discontinuity



$$1/2 \left[\left(\text{Diagram 1} \right) - \left(\text{Diagram 2} \right) \right] = 1/2 \left(\text{Diagram 3} \right)$$

Diagram 1: A green path starts at a blue dot, goes up and to the right, then down and to the right, ending at a green dot. A blue line is below it.

Diagram 2: A green path starts at a blue dot, goes up and to the right, then down and to the right, ending at a green dot. A blue line is below it. A red line is above it.

Diagram 3: A red line starts at a blue dot and goes straight down to a green dot. A blue line is below it.

Summary of potential deformation method

1. Deform the potential so FV is true ground state $T \ll T_{NL}$ (no return flux)
2. Take $T \rightarrow \infty$
 - Picks out $E_{FV}(g)$ $T \gg T_{slosh}$ (only metastable FV decay)
 - Fixes integration contour to be the steepest descent contour passing through the static FV saddle point
3. Deform back

OR

- Compute Z by integrating along the steepest descent contour passing through the static FV saddle point
- OR
- Compute Γ by integrating along the steepest descent contour passing through the bounce, taking the imaginary part, and multiply by 1/2
- Mathematically consistent procedure to get imaginary part out of an analytic real function Z
- Has the right ingredients associated with the necessary limits

Does this procedure give
the decay rate?

3. DIRECT METHOD

Andreassen, Farhi, Frost, MDS (2016)

A direct approach

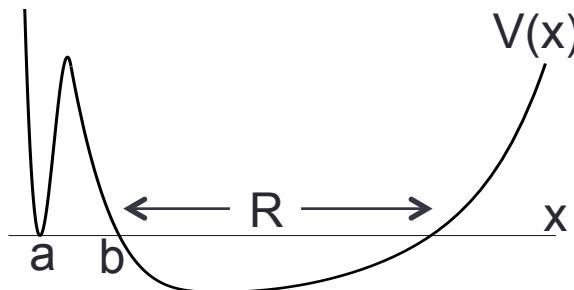
Back to our definition

$$P_{\text{FV}}(T) \equiv \int_{\text{FV}} dx |\psi(x, T)|^2$$

$$\Gamma = - \lim_{\frac{T}{T_{\text{slosh}}} \rightarrow \infty} \lim_{\frac{T}{T_{\text{NL}}} \rightarrow 0} \frac{1}{P_{\text{FV}}} \frac{d}{dT} P_{\text{FV}}$$

$T \gg T_{\text{slosh}}$ (only metastable FV decay)

$T \ll T_{\text{NL}}$ (no return flux)



Propagator from a to x_f in time T

$$D(a, x_f, T) \equiv \int_{x(0)=a}^{x(T)=x_f} \mathcal{D}x e^{iS[x]}$$

- Start with: $\psi(x, t = 0) = \delta(x - a)$
- We will compute

$$\Gamma_R \equiv \lim_{\substack{T/T_{\text{NL}} \rightarrow 0 \\ T/T_{\text{slosh}} \rightarrow \infty}} \frac{1}{P_{\text{FV}}} \frac{dP_R}{dT}$$

$$P_R(T) = \int_R dx_f |D(a, x_f, T)|^2$$

probability of finding ψ
in region R at time T

Step 1: Split up propagator

$$D(a, x_f, T) \equiv \int_{x(0)=a}^{x(T)=x_f} \mathcal{D}x e^{iS[x]}$$

Split path integral into *before b* and *after b*:

$$D(a, x_f, T) = \int_{x(0)=a}^{x(T)=x_f} \mathcal{D}x e^{iS[x]} \int dt \delta(t - t_b[x])$$

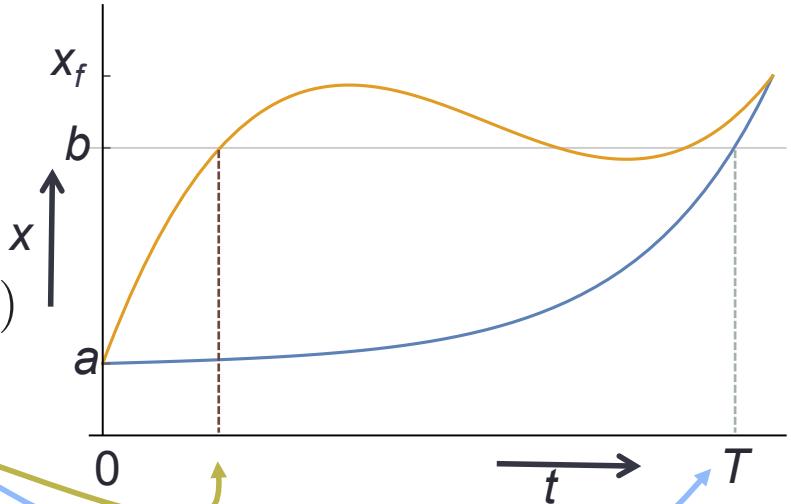
$t_b[x] \equiv \text{First time path } x(t) \text{ hits } b$

$$= \int dt \underbrace{\int_{x(0)=a}^{x(t)=b} \mathcal{D}x e^{iS[x]} \delta(t - t_b[x])}_{\bar{D}(a, b, t)} \underbrace{\int_{x(t)=b}^{x(T)=x_f} \mathcal{D}x e^{iS[x]}}_{D(b, x_f, T-t)}$$

hits b only once, at t

- Regular propagator from b to x_f
- Paths can go back past b

$$D(a, x_f, T) = \int dt \bar{D}(a, b, t) D(b, x_f, T-t)$$



Step 2: Apply $T \ll T_{NL}$

$$D(a, x_f, T) = \int dt \bar{D}(a, b, t) D(b, x_f, T - t)$$

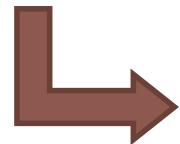
- Hits b only once, at t

- Regular propagator from b to x_f
- Paths can go back past b

$$\begin{aligned} P_R(T) &= \int_R dx_f |D(a, x_f, T)|^2 \\ &= \int dx_f dt_1 dt_2 \bar{D}(a, b, t_1) \bar{D}^*(a, b, t_2) D(b, x_f, T - t_1) D(b, x_f, T - t_2) \\ &\quad \underbrace{\langle x_f, T | b, t_1 \rangle}_{\langle x_f, T | b, t_1 \rangle} \underbrace{\langle b, t_2 | b, t_2 \rangle}_{\langle b, t_2 | b, t_2 \rangle} \end{aligned}$$

$T \ll T_{NL}$ (no return flux)

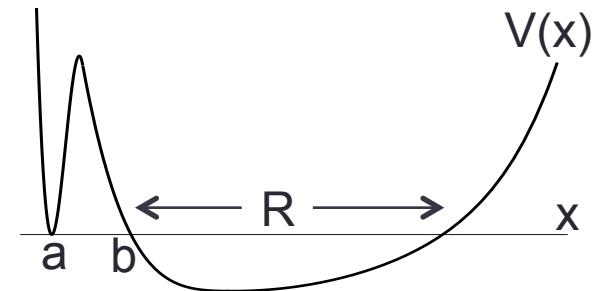
Propagation from b out of R is negligible: $\int_R dx_f |x_f\rangle \langle x_f| = 1$



$$\begin{aligned} P_R(T) &= \int dt_1 dt_2 \bar{D}(a, b, t_1) \bar{D}^*(a, b, t_2) \langle b, t_2 | b, t_1 \rangle \\ &= \int_0^T dt D(a, b, t) \bar{D}^*(a, b, t) + \text{c.c.} \end{aligned}$$

Step 3: Simplify

$$P_R(T) = \int_0^T dt D(a, b, t) \bar{D}^*(a, b, t) + \text{c.c.}$$



$$\Gamma_R \equiv \lim_{\substack{T/T_{\text{NL}} \rightarrow 0 \\ T/T_{\text{slosh}} \rightarrow \infty}} \frac{1}{P_{\text{FV}}} \frac{dP_R}{dT} = \lim_{T \rightarrow \infty} \frac{D(a, b, T) \bar{D}^*(a, b, T)}{\int_{\text{FV}} dx |D(a, x, T)|^2} + \text{c.c.}$$

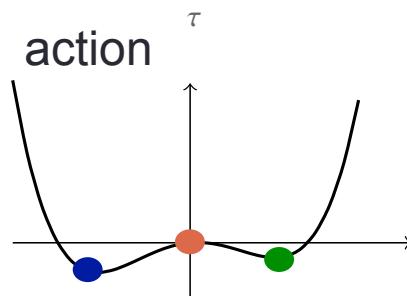
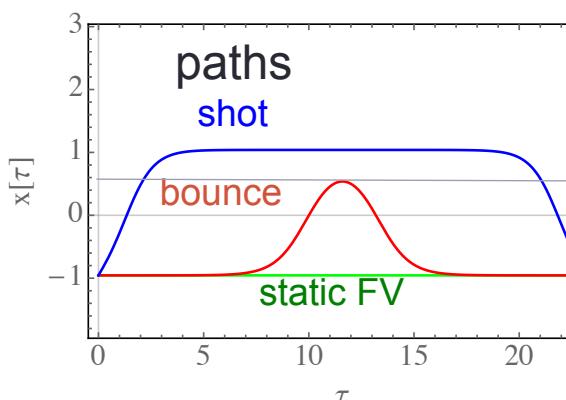
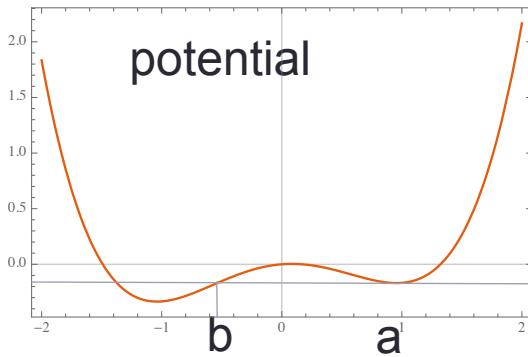
Go to Euclidean time and take

$T \gg T_{\text{slosh}}$

$$\Gamma_R = 2\text{Im} \lim_{T \rightarrow \infty} \left(\frac{\int_{x(-\mathcal{T})=a}^{x(\mathcal{T})=a} \mathcal{D}x e^{-S_E[x]} \delta(\tau_b[x])}{\int_{x(-\mathcal{T})=a}^{x(\mathcal{T})=a} \mathcal{D}x e^{-S_E[x]}} \right)_{\mathcal{T} \rightarrow iT}$$

- Non-perturbative definition of the decay rate
- Does not require analytic continuing potential
- Does not require saddle-point approximation

Expansion



$$\Gamma_R = 2\text{Im} \lim_{T \rightarrow \infty} \left(\frac{\int_{x(-\mathcal{T})=a}^{x(\mathcal{T})=a} \mathcal{D}x e^{-S_E[x]} \delta(\tau_b[x])}{\int_{x(-\mathcal{T})=a}^{x(\mathcal{T})=a} \mathcal{D}x e^{-S_E[x]}} \right)_{\mathcal{T} \rightarrow iT}$$

Bounce dominates numerator

$$\approx \frac{\exp(-S_{\text{shot}}) + \exp(-S_{\text{bounce}})}{\exp(-S_{\text{shot}}) + \exp(-S_{\text{bounce}}) + \exp(-S_{\text{FV}})}$$

FV dominates denominator

$$S_{\text{shot}} = E_{\text{TV}}\mathcal{T} + S_S^0 = iE_{\text{TV}}T + S_S^0$$

$$S_{\text{FV}} = E_{\text{FV}}\mathcal{T} = iE_{\text{FV}}T$$

$$S_{\text{bounce}} = E_{\text{FV}}\mathcal{T} + S_B^0 = iE_{\text{FV}}T + S_B^0$$

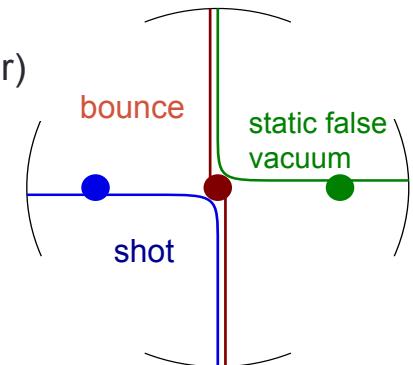
$$S_S^0 > S_B^0 > 0$$

b $=2.5$
 a $=1.2$
 shot must go faster than bounce,
 \rightarrow it has more kinetic energy

$$e^{-S_{\text{FV}}} \gg e^{-S_{\text{bounce}}} \gg e^{-S_{\text{shot}}}$$

Factor of 1/2

$$\Gamma = \text{Im} (\text{FV contour}) = \frac{1}{2} \text{Im} (\text{bounce contour})$$



forces all paths to hit b at time $t=0$

$$\Gamma_R = 2\text{Im} \lim_{T \rightarrow \infty} \left(\frac{\int_{x(-T)=a}^{x(\mathcal{T})=a} \mathcal{D}x e^{-S_E[x]} \delta(\tau_b[x])}{\int_{x(-T)=a}^{x(\mathcal{T})=a} \mathcal{D}x e^{-S_E[x]}} \right)_{\mathcal{T} \rightarrow iT}$$

Expand around bounce: $x(\tau) = \bar{x}(\tau) + \sum \xi_n y_n(\tau)$

Hits b at its maximum

linear in ξ_r

$$\Gamma^{\text{NLO}} = \frac{e^{-S_E[\bar{x}]} e^{-S_E[x_{FV}]}}{e^{-S_E[x_{FV}]}} \lim_{\mathcal{T} \rightarrow \infty} \left| \frac{2 \text{Im} \int d^n \zeta \ J[\tau_*(\zeta), \zeta] e^{-\frac{1}{2} \sum \lambda_i \zeta_i^2}}{\int \mathcal{D} \delta x \ e^{-\frac{1}{2} S_E''[x_{FV}] \delta x^2}} \Theta[\xi_n y_n(0)] \right| \uparrow$$

- Half the fluctuations don't hit b , half do
- Gaussian integral is symmetric.
- Can remove θ -function restriction and multiply by $\frac{1}{2}$

Must hit b

$$\Gamma^{\text{NLO}} = \frac{e^{-S_E[\bar{x}]}}{e^{-S_E[x_{FV}]}} \sqrt{\frac{S_E[\bar{x}]}{2\pi}} \left| \frac{\det'(-\partial_t^2 + V''(\bar{x}(t)))}{\det(-\partial_t^2 + V''(a))} \right|^{-1/2}$$

Agrees with formula from potential deformation method

Summary of tunneling rates

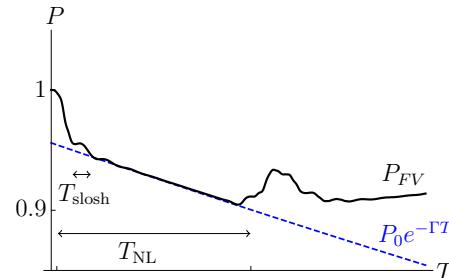
$T \ll T_{NL}$ (no return flux)

Precise definition of decay rate involves **two limits** $\Gamma = -\lim_{\frac{T}{T_{\text{sho}}} \rightarrow \infty} \lim_{\frac{T}{T_{NL}} \rightarrow 0} \frac{1}{P_{FV}} \frac{d}{dT} P_{FV}$

$T \gg T_{\text{sho}}$ (remove transients)

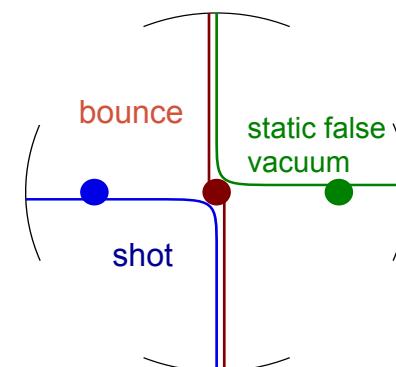
Three methods to compute Γ

1. **Solve Schrodinger's equation**
- Impractical for QFT



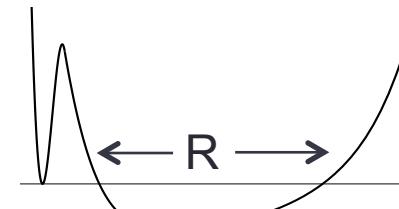
2. **Deform potential** to stabilize false vacuum
- Take $T \rightarrow \infty$ limit
- Deform back and compute imaginary part

Is the result the decay rate?



3. **Direct approach** using Minkowski space causal propagators
- Does not rely on saddle-point approximation
- Does not rely on deforming potential
- QFT derivation is simple – no bold leap of faith
- Non-perturbative formula

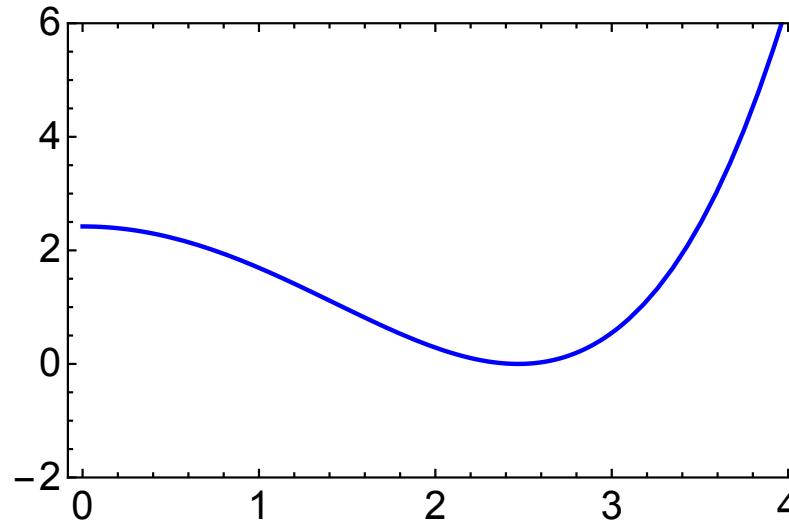
$$\Gamma_R = 2\text{Im} \lim_{T \rightarrow \infty} \left(\frac{\int_{x(-\mathcal{T})=a}^{x(\mathcal{T})=a} \mathcal{D}x e^{-S_E[x]} \delta(\tau_b[x])}{\int_{x(-\mathcal{T})=a}^{x(\mathcal{T})=a} \mathcal{D}x e^{-S_E[x]}} \right)_{\mathcal{T} \rightarrow iT}$$



EFFECTIVE POTENTIALS AND DECAY RATES

COMPUTING EFFECTIVE POTENTIALS

How do we compute V_{eff} ?



Classical potential: $V(h) = \Lambda + m^2 h^2 + \lambda h^4$

- Renormalizable
- Three parameters (Λ , m and λ), measured from data

How can the quantum-corrected potential be computed?

How do we compute V_{eff} ?

Method 1:

$$\int \mathcal{D}H e^{i\Gamma} \equiv \int \mathcal{D}H \underbrace{\mathcal{D}\psi \cdots \mathcal{D}A}_{\text{Integrate out everything but } H} e^{iS}$$

Effective Action

Classical action

$$\Gamma = \int d^4x \left\{ -Z[H]H\Box H - V_{\text{eff}}(H) + \dots \right\}$$

Problems:

- Generally non-local (has nasty things like $\ln \frac{1 + \Box/m_t^2}{H^2}$ in it)
- Nearly impossible to compute
- Can't include loops of H itself this way

OK if $H \approx \langle H \rangle$

If we integrate over everything,
effective action is just a number

$$e^{i\Gamma} = \int \mathcal{D}H \cdots \mathcal{D}A e^{iS}$$

Method 2: Legendre transform

Classical action

$$\frac{\delta S}{\delta H} \Big|_{H=v} = 0$$

↑
Classical minimum

We want an effective action

$$\frac{\delta \Gamma}{\delta H} \Big|_{H=H_q} = 0$$

↑
True quantum minimum

1. Compute $W[J]$ $e^{W[J]} \equiv \int \mathcal{D}H \dots \mathcal{D}A e^{i \int d^4x \{ \mathcal{L} + JH \}}$

2. Solve $H = \frac{\partial W}{\partial J}$ for $J[H]$

3. Compute $\Gamma[H] = W[J[H]] - \int d^4x H J[H]$

Current introduced by hand
So that Γ depends on something

Has the property that $\frac{\delta \Gamma}{\delta H} = J[H]$ so that $\frac{\delta \Gamma}{\delta H} = 0$ when $J=0$ (i.e. in original theory)

- Agrees with method 1 in perturbation theory

What do you get?

Tree-level (classical)

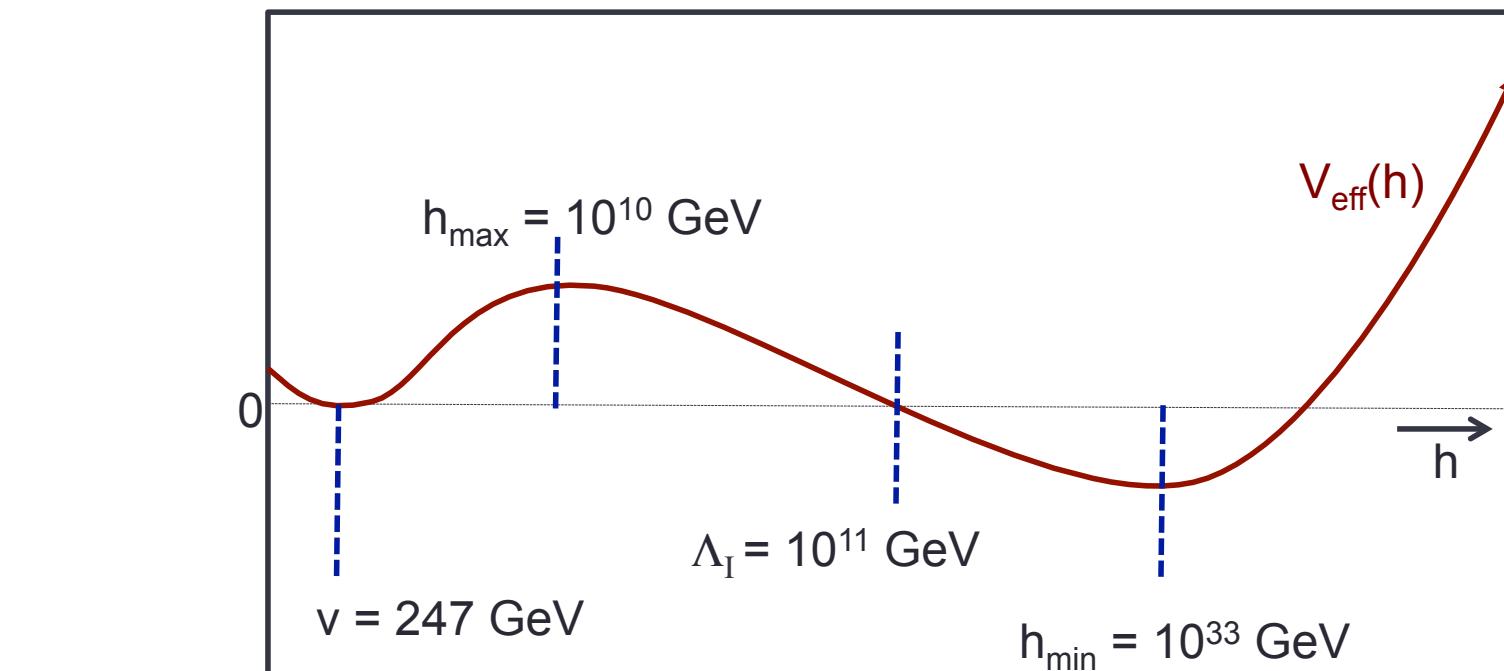
$$V_{\text{eff}} = \frac{1}{4} \lambda h^4 - m^2 h^2$$

$$+ h^4 \frac{1}{2048\pi^2} \left[-5g_1^4 + 6(g_1^2 + g_2^2)^2 \ln \frac{h^2(g_1^2 + g_2^2)}{4\mu^2} - 10g_1^2 g_2^2 - 15g_2^4 + 12g_2^4 \ln \frac{g_2^2 h^2}{4\mu^2} + 144y_t^4 - 96y_t^4 \ln \frac{y_t^2 h^2}{2\mu^2} \right]$$

$$\frac{-1}{256\pi^2} \left[\xi_B g_1^2 \left(\ln \frac{\lambda h^4 (\xi_B g_1^2 + \xi_W g_2^2)}{4\mu^4} - 3 \right) + \xi_W g_2^2 \left(\ln \frac{\lambda^3 h^{12} \xi_W^2 g_2^4 (\xi_B g_1^2 + \xi_W g_2^2)}{64\mu^{12}} - 9 \right) \right] \lambda h^4$$

one-loop

$+ \dots$



What do you get?

Tree-level (classical)

$$\begin{aligned}
 V_{\text{eff}} = & \frac{1}{4} \lambda h^4 - m^2 h^2 \\
 & + h^4 \frac{1}{2048\pi^2} \left[-5g_1^4 + 6(g_1^2 + g_2^2)^2 \ln \frac{h^2(g_1^2 + g_2^2)}{4\mu^2} - 10g_1^2 g_2^2 - 15g_2^4 + 12g_2^4 \ln \frac{g_2^2 h^2}{4\mu^2} + 144y_t^4 - 96y_t^4 \ln \frac{y_t^2 h^2}{2\mu^2} \right] \\
 & + \frac{-1}{256\pi^2} \left[\xi_B g_1^2 \left(\ln \frac{\lambda h^4 (\xi_B g_1^2 + \xi_W g_2^2)}{4\mu^4} - 3 \right) + \xi_W g_2^2 \left(\ln \frac{\lambda^3 h^{12} \xi_W^2 g_2^4 (\xi_B g_1^2 + \xi_W g_2^2)}{64\mu^{12}} - 9 \right) \right] \lambda h^4 \\
 & + \dots
 \end{aligned}$$

one-loop

Two curious features

- 1. Not gauge-invariant
- 2. Large logarithms

1. Gauge-dependence

Method 1 to compute Γ **is** gauge-invariant:

$$\int \mathcal{D}H e^{i\Gamma} \equiv \int \mathcal{D}H \underbrace{\mathcal{D}\psi \cdots \mathcal{D}A e^{iS}}_{}$$

Completely integrate over gauge-orbits

Action/energy at minimum also gauge-invariant: $e^{i\Gamma} = \int \mathcal{D}H \cdots \mathcal{D}A e^{iS}$

Method 2 to compute Γ introduces a **charged source J**

$$e^{W[J]} \equiv \int \mathcal{D}H \cdots \mathcal{D}A e^{i \int d^4x \{ \mathcal{L} + JH \}}$$

$$\Gamma = W - HJ$$

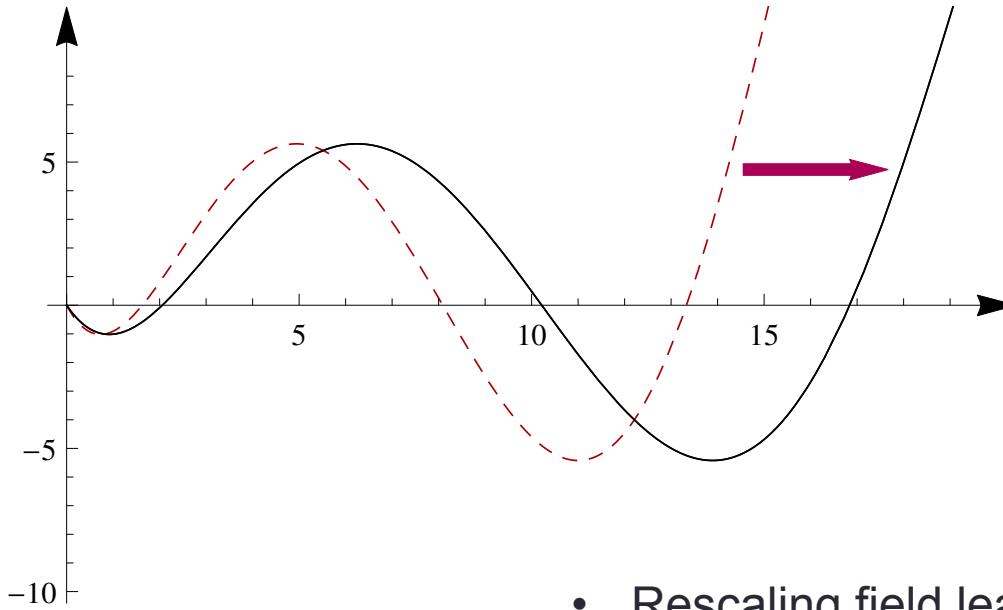
$$\frac{\delta \Gamma}{\delta H} = J$$

- Action **away from minimum** has **current** present
- Action **at minimum** has **no current**, should be gauge-invariant

Encoded in
Nielsen identity

$$\left[\xi \frac{\partial}{\partial \xi} + C(h, \xi) \frac{\partial}{\partial h} \right] V_{\text{eff}}(h, \xi) = 0$$

Potential at minimum indep. of rescaling

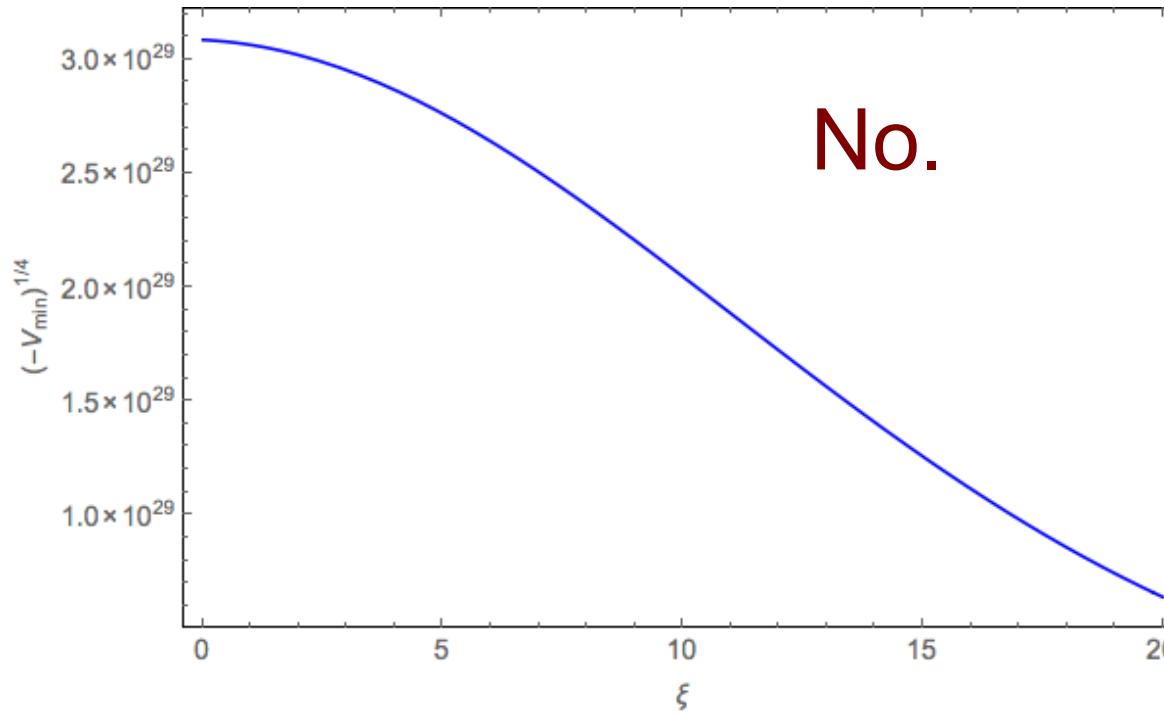


- Rescaling field leaves V_{\min} unchanged

Nielsen identity

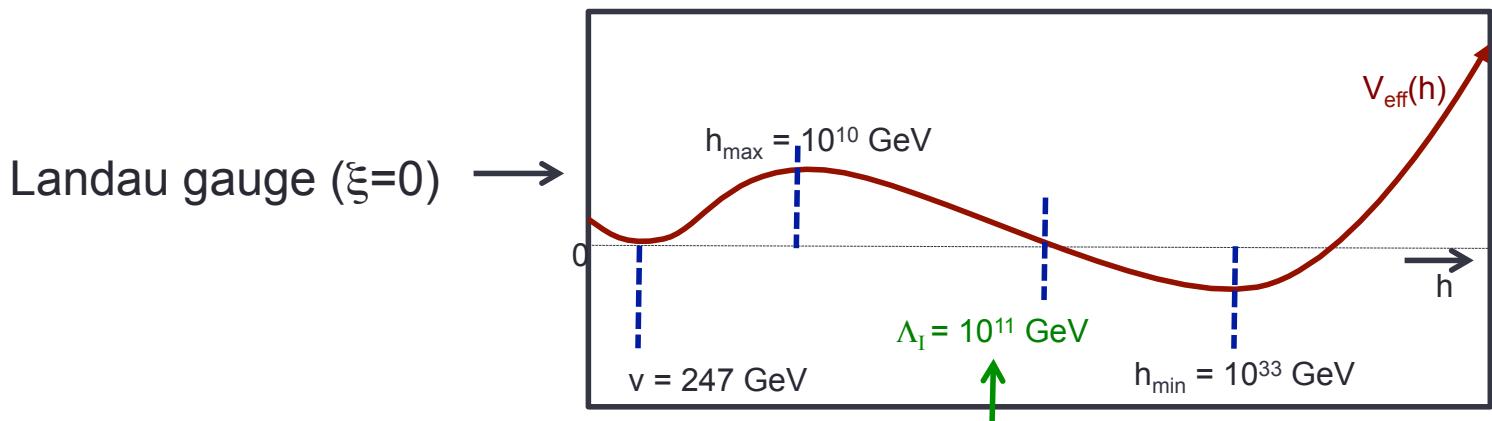
$$\left[\xi \frac{\partial}{\partial \xi} + C(h, \xi) \frac{\partial}{\partial h} \right] V_{\text{eff}}(h, \xi) = 0$$

But is it?

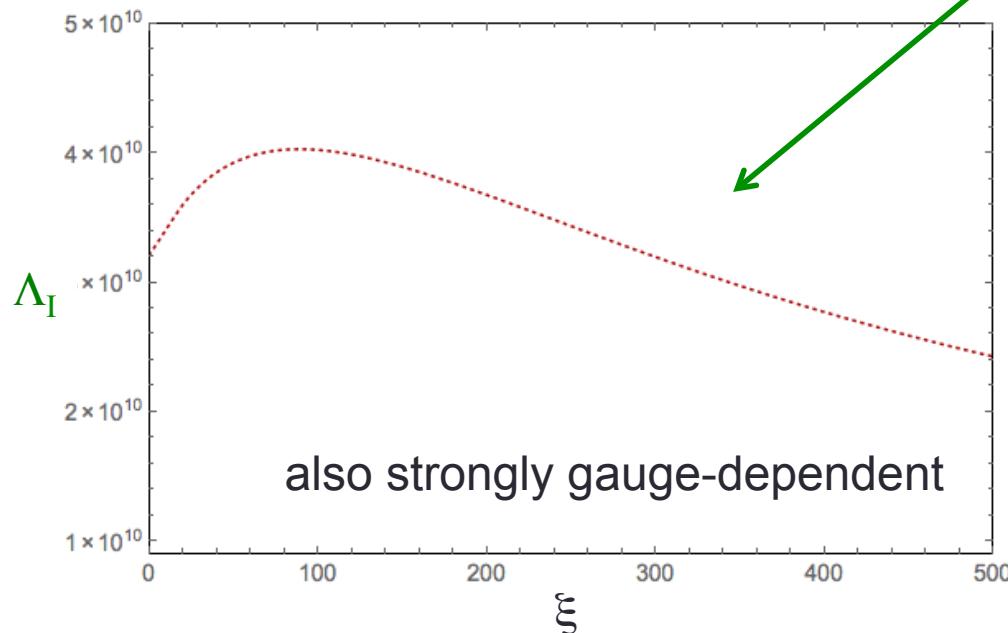


$(-V_{\min})^{1/4}$ appears linearly-dependent on gauge parameter ξ

What about field values?



Instability scale $\Lambda_I = \text{value of } h \text{ where } V(h) = 0$



- h_{min} also gauge dependent
- h_{max} also gauge dependent
- ...

2. Large Logarithms

Can be resummed with RGE:

Explicit μ dependence

$$\left(\mu \frac{\partial}{\partial \mu} + \beta_i \frac{\partial}{\partial g_i} - \gamma h \frac{\partial}{\partial h} \right) V_{\text{eff}} = 0$$

compensated for by rescaling couplings and fields

- Same RGE as 1PI Green's functions or off-shell matrix elements
- Observables/S-matrix elements satisfy simpler RGE:

$$\left(\mu \frac{\partial}{\mu} + \beta_i \frac{\partial}{\partial g_i} \right) \sigma = 0$$

- Field-rescaling term canceled by LSZ wavefunction Z-factors



Effective potential depends on the normalization of fields??!!

Resum logarithms

1. Compute V_{eff} to fixed order (say 2-loops) at scale (say) $\mu_0 \sim 100 \text{ GeV}$
2. Solve RGE $\left(\mu \frac{\partial}{\partial \mu} + \beta_i \frac{\partial}{\partial g_i} - \gamma h \frac{\partial}{\partial h} \right) V_{\text{eff}} = 0$

$$V_{\text{eff}}(h, g_i, \mu) \rightarrow V_{\text{eff}}(e^{\Gamma(\mu_0, \mu)} h, g_i(\mu), \mu)$$

$\Gamma(\mu_0, \mu) \equiv \int_{\mu_0}^{\mu} \gamma(\mu') d \ln \mu'$

3. Set $\mu \sim h$

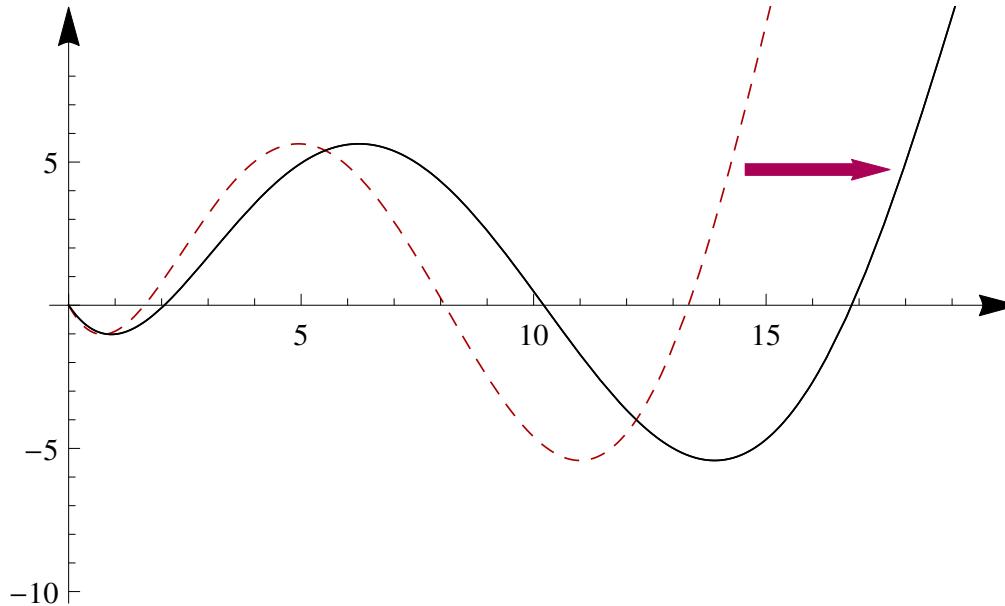
$$V_{\text{eff}}(h, \mu_0) = V_{\text{eff}}(e^{\Gamma(\mu_0, h)} h, g_i(h), h)$$

Potential depends on scale μ_0 where it is calculated??!!



$$\left(\frac{\partial}{\partial \mu_0} - \gamma h \frac{\partial}{\partial h} \right) V(h, \mu_0) = 0$$

Potential at minimum



Nielsen identity (gauge invariance)

$$\left[\xi \frac{\partial}{\partial \xi} + C(h, \xi) \frac{\partial}{\partial h} \right] V_{\text{eff}}(h, \xi) = 0$$

Calculation-scale invariance

$$\left(\frac{\partial}{\partial \mu_0} - \gamma h \frac{\partial}{\partial h} \right) V(h, \mu_0) = 0$$

V_{min} should be gauge invariant and independent of how it is calculated

Even gauge-invariant Γ is unphysical

Even if we source a gauge-invariant field $e^{W[J]} \equiv \int \mathcal{D}H \dots \mathcal{D}A e^{i \int d^4x \{\mathcal{L} + JH\}}$

$$\left. \begin{array}{l} e^{W[J]} \equiv \int \mathcal{D}H \dots \mathcal{D}A e^{i \int d^4x \{\mathcal{L} + JH^\dagger H\}} \\ e^{W[J]} \equiv \int \mathcal{D}H \dots \mathcal{D}A e^{i \int d^4x \{\mathcal{L} + J|H|\}} \end{array} \right\} \quad \Gamma(h) \text{ is now gauge-invariant}$$

Effective potential still depends on how it is calculated $\left(\frac{\partial}{\partial \mu_0} - \gamma h \frac{\partial}{\partial h} \right) V(h, \mu_0) = 0$

- This is OK.
- Off-shell quantities can be unphysical
- **Observables should be physical**

- S-matrix elements
- Vacuum energy (V_{\min})
- Tunnelling rates
- Critical temperature

But are they?

What about field values?

- Instability scale?
- Inflation scale?
- Planck/new physics sensitivity?

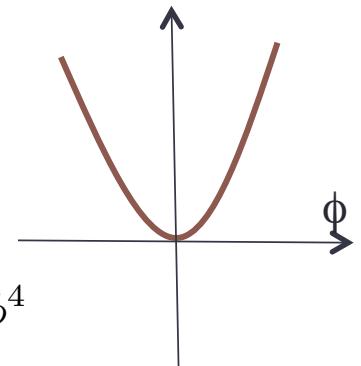
Are these questions about observables?

SCALAR QED

Scalar QED

$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}^2 + \frac{1}{2}|D_\mu\phi|^2 - V(\phi)$$

$$\uparrow \sim V_0(\phi) = \frac{\lambda}{24}\phi^4$$



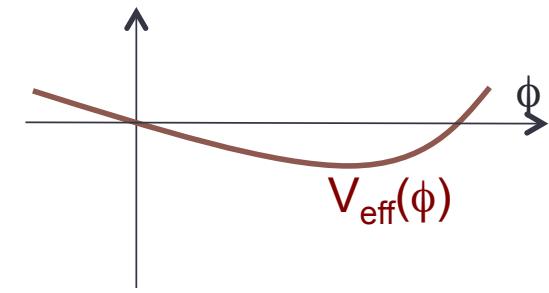
- mass term gives small corrections, so we drop it

1-loop potential in R_ξ gauges:

$$V_1(\phi) = \phi^4 \frac{\hbar}{16\pi^2} \left[\frac{3}{4}e^4 \left(\ln \frac{e^2\phi^2}{\mu^2} - \frac{5}{6} \right) + \frac{\lambda^2}{16} \left(\ln \frac{\lambda\phi^2}{2\mu^2} - \frac{3}{2} \right) \right. \\ \left. + \left(\frac{\lambda^2}{144} - \frac{1}{12}e^2\lambda\xi \right) \left(\ln \frac{\phi^2}{\mu^2} - \frac{3}{2} \right) + \frac{1}{4}K_+^4 \ln K_+^2 + \frac{1}{4}K_-^4 \ln K_-^2 \right] \\ K_\pm^2 = \frac{1}{12} \left(\lambda \pm \sqrt{\lambda^2 - 24\lambda e^2 \xi} \right)$$

- Not gauge-invariant

- For most values of e and λ , there is no minimum
- When $\lambda \approx \frac{e^4}{16\pi^2}$ $\Rightarrow V_0 \approx V_1$ 
- And.... V_{\min} depends on ξ

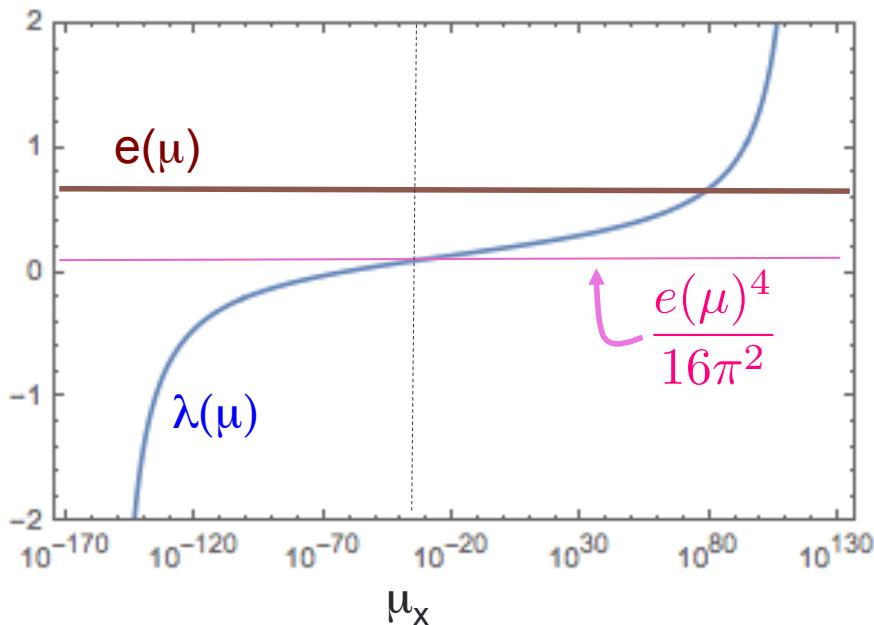


Spontaneous symmetry breaking

When is $\lambda \approx \frac{e^4}{16\pi^2}$?

Solve
RGEs:

$$\left. \begin{aligned} \beta_e &= \frac{\hbar}{16\pi^2} \left(\frac{e^3}{3} \right) + \dots \\ \beta_\lambda &= \frac{\hbar}{16\pi^2} \left(36e^4 - 12e^2\lambda + \frac{10\lambda^2}{3} \right) \end{aligned} \right\} \quad \begin{aligned} e^2(\mu) &= \frac{e^2(\mu_0)}{1 - \frac{e^2(\mu_0)}{24\pi^2} \ln \frac{\mu}{\mu_0}} \\ \lambda(\mu) &= \frac{e^2(\mu)}{10} \left[19 + \sqrt{719} \tan \left(\frac{\sqrt{719}}{2} \ln \frac{e(\mu)^2}{C} \right) \right] \end{aligned}$$



- e runs relatively slowly
- For any e , λ runs through all values
- There is always a scale μ_x where

$$\lambda(\mu_x) \approx \frac{e(\mu_x)^4}{16\pi^2}$$

- Near this scale, V_{eff} is perturbative

Proper loop expansion

$$V_0(\phi) = \frac{\lambda}{24}\phi^4$$

$$V_1(\phi) = \phi^4 \frac{\hbar}{16\pi^2} \left[\frac{3}{4}e^4 \left(\ln \frac{e^2\phi^2}{\mu^2} - \frac{5}{6} \right) + \frac{\lambda^2}{16} \left(\ln \frac{\lambda\phi^2}{2\mu^2} - \frac{3}{2} \right) \right. \\ \left. + \left(\frac{\lambda^2}{144} - \frac{1}{12}e^2\lambda\xi \right) \left(\ln \frac{\phi^2}{\mu^2} - \frac{3}{2} \right) + \frac{1}{4}K_+^4 \ln K_+^2 + \frac{1}{4}K_-^4 \ln K_-^2 \right]$$

$\uparrow \quad \uparrow$

$$K_{\pm}^2 = \frac{1}{12} \left(\lambda \pm \sqrt{\lambda^2 - 24\lambda e^2 \xi} \right)$$

Comparable when

$$\lambda \approx \hbar \frac{e^4}{16\pi^2}$$

- Then V_0 and V_1 of order \hbar

These terms all have extra \hbar suppression

Expanding in \hbar with $\lambda \sim \hbar$

order \hbar : $V^{\text{LO}} = \frac{\lambda}{24}\phi^4 + \frac{\hbar e^4}{16\pi^2}\phi^4 \left(-\frac{5}{8} + \frac{3}{2} \ln \frac{e\phi}{\mu} \right) \longrightarrow V_{\text{min}}^{\text{LO}} = -\frac{3}{128\pi^2}e^4 \langle \phi \rangle^4$

order \hbar^2 : $V^{\text{NLO}} = \frac{\hbar e^2 \lambda}{16\pi^2} \phi^4 \left(\frac{\xi}{8} - \frac{\xi}{24} \ln \frac{e^2 \lambda \xi \phi^4}{6\mu^4} \right) \longrightarrow V_{\text{min}}^{\text{NLO}} = \dots$

Problem: higher-loop contributions also of order \hbar^2

2-Loop potential in scalar QED

- Known in Landau gauge
- Some terms computed by Kang (1974), not in $\overline{\text{MS}}$
- Some terms at order $e^6 \hbar^2$ unknown

We computed all the relevant 2-loop graphs:

$$\text{Diagram 1} = \frac{\hbar^2 \phi^4 e^6}{(16\pi^2)^2} \xi \left[-12 \ln^2 \frac{e\phi}{\mu} + \left(8 - 3 \ln \frac{\lambda\xi}{6e^2} \right) \ln \frac{e\phi}{\mu} - \frac{5}{2} - \frac{\pi^2}{16} - \frac{3}{16} \ln^2 \frac{\lambda\xi}{6e^2} + \ln \frac{\lambda\xi}{6e^2} \right]$$

$$\text{Diagram 2} = \frac{\hbar^2 \phi^4 e^6}{(16\pi^2)^2} \left[(2 + 6\xi) \ln^2 \frac{e\phi}{\mu} - (3 + 7\xi) \ln \frac{e\phi}{\mu} + \frac{7}{4} + \frac{\pi^2}{8} + \frac{15}{4}\xi + \frac{3\pi^2}{8}\xi \right]$$

$$\text{Diagram 3} = \frac{\hbar^2 \phi^4 e^6}{(16\pi^2)^2} \left[(18 + 6\xi) \ln^2 \frac{e\phi}{\mu} - (21 + 7\xi) \ln \frac{e\phi}{\mu} + \frac{47}{4} + \frac{7\pi^2}{24} + \frac{15}{4}\xi + \frac{3\pi^2}{8}\xi \right]$$

$$\text{Diagram 4} = \frac{\hbar^2 \phi^4 e^6}{(16\pi^2)^2} \xi \left[-12 \ln^2 \frac{e\phi}{\mu} + 14 \ln \frac{e\phi}{\mu} - \frac{15}{2} - \frac{3\pi^2}{4} \right]$$

Then the relevant part of the 2-loop potential is

$$V_2 = \left(\frac{\hbar}{16\pi^2} \right)^2 e^6 \phi^4 \left[(10 - 6\xi) \ln^2 \frac{e\phi}{\mu} + \left(-\frac{62}{3} + 4\xi - \frac{3}{2}\xi \ln \frac{\lambda\xi}{6e^2} \right) \ln \frac{e\phi}{\mu} + \xi \left(-\frac{1}{2} + \frac{1}{4} \ln \frac{\lambda\xi}{6e^2} \right) + \frac{71}{6} \right] + \dots \text{ terms of order } \hbar^3$$

Potential at minimum

$$V^{\text{LO}} = \frac{\lambda}{24}\phi^4 + \frac{\hbar e^4}{16\pi^2}\phi^4 \left(-\frac{5}{8} + \frac{3}{2} \ln \frac{e\phi}{\mu} \right)$$

$$V^{\text{NLO}} = \frac{\hbar e^2 \lambda}{16\pi^2} \phi^4 \left(\frac{\xi}{8} - \frac{\xi}{24} \ln \frac{e^2 \lambda \xi \phi^4}{6\mu^4} \right) + \frac{\hbar^2 e^6}{(16\pi^2)^2} \phi^4 \left[(10 - 6\xi) \ln^2 \frac{e\phi}{\mu} + \left(-\frac{62}{3} + 4\xi - \frac{3}{2}\xi \ln \frac{\lambda\xi}{6e^2} \right) \ln \frac{e\phi}{\mu} + \xi \left(-\frac{1}{2} + \frac{1}{4} \ln \frac{\lambda\xi}{6e^2} \right) + \frac{71}{6} \right]$$

- Solve $V'(\phi=v) = 0$ for $\lambda(v)$:

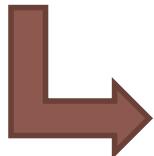
$$\lambda = \frac{\hbar e^4}{16\pi^2} \left(6 - 36 \ln \frac{ev}{\mu} \right) + \frac{\hbar e^6}{(16\pi^2)^2} \left\{ -160 - 24\xi + (376 + 90\xi) \ln \frac{ev}{\mu} - 240 \ln^2 \frac{ev}{\mu} + 9\xi \ln \left[\frac{\xi \hbar \mu^2}{16\pi^2 v^2} \left(1 - 6 \ln \frac{ev}{\mu} \right) \right] \right\}$$

- Plug in to $V(v)$:

$$V_{\min} = v^4 \frac{e^4 \hbar}{16\pi^2} \left(-\frac{3}{8} \right) + v^4 \frac{e^6 \hbar^2}{(16\pi^2)^2} \frac{1}{12} \left\{ 62 - 9\xi + (-60 + 18\xi) \ln \frac{ev}{\mu} + \frac{9}{2}\xi \ln \left[\frac{e^2 \xi \hbar}{16\pi^2} \left(1 - 6 \ln \frac{ev}{\mu} \right) \right] \right\}$$

Still gauge-dependent!

Problem: $v = \langle \phi \rangle$ is gauge-dependent



Express V_{\min} in terms of only other dimensionful scale: μ

In terms of μ_X

Define μ_X by

$$\lambda(\mu_X) \equiv \frac{\hbar}{16\pi^2} e^4(\mu_X) \left[6 - 36 \ln[e(\mu_X)] \right]$$

- Tree-level vev is $v=\mu_X$
- Exact (non-perturbative) definition of μ_X

Then, vev is:

$$v = \mu_X + \mu_X \frac{\hbar e^2}{16\pi^2} \left\{ -\frac{40}{9} + \frac{94}{9} \ln e - \frac{20}{3} \ln^2 e - \frac{\xi}{2} + \frac{3}{2} \xi \ln e + \frac{\xi}{4} \ln \left[\frac{\xi \hbar}{16\pi^2} (1 - 6 \ln e) \right] - \frac{1}{6} \xi + \xi \ln e \right\}.$$

- gauge-dependent vev is OK – not physical

Potential at minimum is:

$$V_{\min} = \frac{e^4 \hbar}{16\pi^2} \mu_X^4 \left(-\frac{3}{8} \right) + \frac{e^6 \hbar}{(16\pi^2)^2} \mu_X^4 \left(\frac{71}{6} - \frac{62}{3} + 10 \ln^2 e \right) + \frac{e^6 \hbar}{(16\pi^2)^2} \mu_X^4 \left(\frac{\xi}{4} - \frac{3}{2} \xi \ln e \right)$$

- gauge-dependent vacuum energy is **not OK**

Still gauge-dependent!

What's missing?

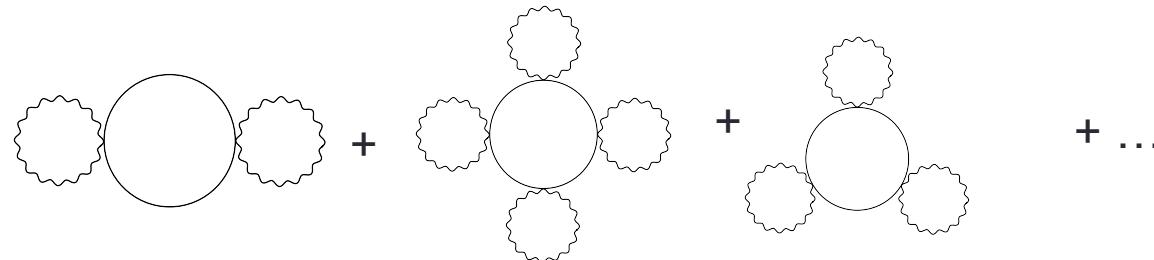
More diagrams!

Daisy resummation

Higher order graphs can scale like inverse powers of λ :

$$\propto (e^2)^3 (e^2 \phi^2)^3 \int \frac{d^4 k}{2\pi^4} \left(\frac{i}{k^2 - \frac{\lambda}{2} \phi^2} \right)^3 \propto \phi^4 \frac{e^{12}}{\lambda}$$

Only one series of graphs contribute at order $\sim \hbar^2$



“daisy diagrams”

We can sum the series:

$$V^{e^6 \text{daisies}} = \phi^4 \frac{\hbar}{16\pi^2} \left(-\frac{e^2 \lambda \xi}{24} \right) \left[\frac{\hat{\lambda}(\phi)}{\lambda} + \left(1 - \frac{\hat{\lambda}(\phi)}{\lambda} \right) \ln \left(1 - \frac{\hat{\lambda}(\phi)}{\lambda} \right) \right]$$

$$\hat{\lambda}(\phi) \equiv \frac{\hbar e^4}{16\pi^2} \left(6 - 36 \ln \frac{e\phi}{\mu} \right)$$

Full potential at NLO:

$$\begin{aligned}
 V^{\text{NLO}} = & \frac{\hbar e^2 \lambda}{16\pi^2} \phi^4 \left(\frac{\xi}{8} - \frac{\xi}{24} \ln \frac{e^2 \lambda \xi \phi^4}{6\mu^4} \right) \\
 & + \frac{\hbar^2 e^6}{(16\pi^2)^2} \phi^4 \left[(10 - 6\xi) \ln^2 \frac{e\phi}{\mu} + \left(-\frac{62}{3} + 4\xi - \frac{3}{2}\xi \ln \frac{\lambda\xi}{6e^2} \right) \ln \frac{e\phi}{\mu} + \xi \left(-\frac{1}{2} + \frac{1}{4} \ln \frac{\lambda\xi}{6e^2} \right) + \frac{71}{6} \right] \\
 & + \phi^4 \frac{\hbar e^2 \lambda}{16\pi^2} \left(-\frac{\xi}{24} \right) \left[\frac{\widehat{\lambda}(\phi)}{\lambda} + \left(1 - \frac{\widehat{\lambda}(\phi)}{\lambda} \right) \ln \left(1 - \frac{\widehat{\lambda}(\phi)}{\lambda} \right) \right]
 \end{aligned}$$

Now... vacuum energy is gauge-invariant!

$$V_{\min} = -\frac{3\hbar e^4}{128\pi^2} \mu_X^4 + \frac{e^6 \hbar^2}{(16\pi^2)^2} \mu_X^4 \left(\frac{71}{6} - \frac{62}{3} \ln e + 10 \ln^2 e \right)$$

Field values are still gauge-dependent:

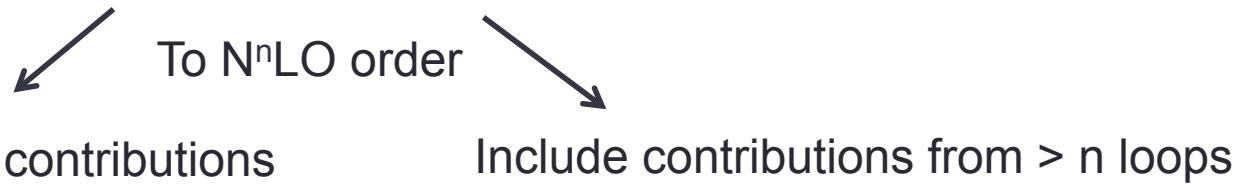
$$v = \mu_X + \mu_X \frac{\hbar e^2}{16\pi^2} \left\{ -\frac{40}{9} + \frac{94}{9} \ln e - \frac{20}{3} \ln^2 e - \frac{\xi}{2} + \frac{3}{2} \xi \ln e + \frac{\xi}{4} \ln \left[\frac{\xi \hbar}{16\pi^2} (1 - 6 \ln e) \right] - \frac{1}{6} \xi + \xi \ln e \right\}.$$

$$\Lambda_I = \mu_I + \mu_I \frac{\hbar e^2}{16\pi^2} \left\{ -\frac{77}{9} + \frac{124}{9} \ln e - \frac{20}{3} \ln^2 e - \frac{\xi}{2} + \frac{3}{2} \xi \ln e + \frac{\xi}{4} \ln \left[\frac{\xi \hbar}{16\pi^2} (1 - 6 \ln e) \right] - \frac{5}{12} \xi + \xi \ln e \right\}.$$

STANDARD MODEL

Lessons from scalar QED

1. Gauge invariance requires consistent expansion in \hbar



2. Don't resum logs by solving RGE for V_{eff}

$$\left(\mu \frac{\partial}{\partial \mu} + \beta_i \frac{\partial}{\partial g_i} - \gamma h \frac{\partial}{\partial h} \right) V_{\text{eff}} = 0$$

- Mixes up orders in \hbar in an uncontrolled way

3. Do resum logs by using couplings at some scale μ_X

- Natural condition for μ_X is that $V_{\text{LO}}'(\phi=\mu_X) = 0$

4. Don't express V_{min} in terms of $v = \langle \phi \rangle$

- Express V_{min} in terms of μ_X instead

Standard Model

$$V^{(\text{LO})}(h) = \frac{1}{4}\lambda h^4 + h^4 \frac{1}{2048\pi^2} \left[-5g_1^4 + 6(g_1^2 + g_2^2)^2 \ln \frac{h^2(g_1^2 + g_2^2)}{4\mu^2} \right. \\ \left. - 10g_1^2 g_2^2 - 15g_2^4 + 12g_2^4 \ln \frac{g_2^2 h^2}{4\mu^2} + 144y_t^4 - 96y_t^4 \ln \frac{y_t^2 h^2}{2\mu^2} \right]$$

Tree-level

Part of 1-loop $\lambda \sim \mathcal{O}(\hbar)$

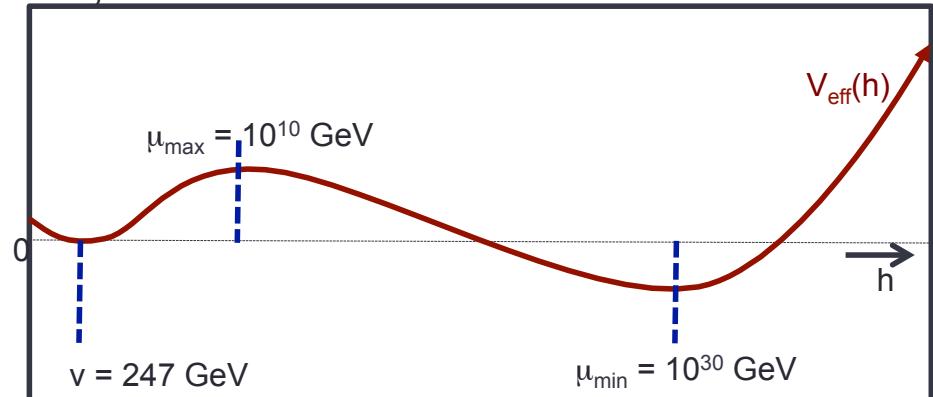
- Scale $h = \mu_X$ where $\frac{d}{dh}V^{(\text{LO})}(h) = 0$ is

$$\lambda = \frac{1}{256\pi^2} \left[g_1^4 + 2g_1^2 g_2^2 + 3g_2^4 - 48y_t^4 - 3(g_1^2 + g_2^2)^2 \ln \frac{g_1^2 + g_2^2}{4} - 6g_2^4 \ln \frac{g_2^2}{4} + 48y_t^4 \ln \frac{y_t^2}{2} \right]$$

- Run couplings with 3-loop β -functions, find numerical solutions

$$\mu_X^{\max} = 2.46 \times 10^{10} \text{ GeV}$$

$$\mu_X^{\min} = 3.43 \times 10^{30} \text{ GeV}$$



Standard Model at NLO

- We know the 1-loop contribution to V_{NLO}

$$V^{(1,\text{NLO})}(h) = \frac{-1}{256\pi^2} \left[\xi_B g_1^2 \left(\ln \frac{\lambda h^4 (\xi_B g_1^2 + \xi_W g_2^2)}{4\mu^4} - 3 \right) + \xi_W g_2^2 \left(\ln \frac{\lambda^3 h^{12} \xi_W^2 g_2^4 (\xi_B g_1^2 + \xi_W g_2^2)}{64\mu^{12}} - 9 \right) \right] \lambda h^4$$

- We know the 2-loop contribution to V_{NLO} in Landau gauge

$$\begin{aligned} \lambda_{\text{eff}}^{(2)} = & \frac{1}{(4\pi)^4} \left[8g_3^2 y_t^4 (3r_t^2 - 8r_t + 9) + \frac{1}{2} y_t^6 (-6r_t r_W - 3r_t^2 + 48r_t - 6r_{tW} - 69 - \pi^2) + \right. \\ & + \frac{3y_t^2 g_Y^4}{16} (8r_W + 4r_Z - 3r_t^2 - 6r_t r_Z - 12r_t + 12r_{tW} + 15 + 2\pi^2) + \\ & + \frac{y_t^2 g_Y^4}{48} (27r_t^2 - 54r_t r_Z - 68r_t - 28r_Z + 189) + \frac{y_t^2 g_2^2 g_Y^2}{8} (9r_t^2 - 18r_t r_Z + 4r_t + 44r_Z - 57) + \\ & + \frac{g_2^6}{192} (36r_t r_Z + 54r_t^2 - 414r_W r_Z + 69r_W^2 + 1264r_W + 156r_Z^2 + 632r_Z - 144r_{tW} - 2067 + 90\pi^2) + \\ & \left. + \frac{g_2^4 g_Y^2}{192} (12r_t r_Z - 6r_t^2 - 6r_W (53r_Z + 50) + 213r_W^2 + 4r_Z (57r_Z - 91) + 817 + 46\pi^2) + \right. \\ & \left. + \frac{g_2^2 g_Y^4}{576} (132r_t r_Z - 66r_t^2 + 306r_W r_Z - 153r_W^2 - 36r_W + 924r_Z^2 - 4080r_Z + 4359 + \right. \\ & \left. + \frac{g_Y^6}{576} (6r_Z (34r_t + 3r_W - 470) - 102r_t^2 - 9r_W^2 + 708r_Z^2 + 2883 + 206\pi^2) + \right. \\ & \left. + \frac{y_t^4}{6} (4g_Y^2 (3r_t^2 - 8r_t + 9) - 9g_2^2 (r_t - r_W + 1)) + \frac{3}{4} (g_2^6 - 3g_2^4 y_t^2 + 4y_t^6) \text{Li}_2 \frac{g_2^2}{2y_t^2} + \right. \\ & \left. + \frac{y_t^2}{48} \xi \left(\frac{g_2^2 + g_Y^2}{2y_t^2} \right) \left(9g_2^4 - 6g_2^2 g_Y^2 + 17g_Y^4 + 2y_t^2 (7g_Y^2 - 73g_2^2 + \frac{64g_2^4}{g_Y^2 + g_2^2}) \right) + \right. \\ & \left. + \frac{g_2^2}{64} \xi \left(\frac{g_2^2 + g_Y^2}{g_2^2} \right) \left(18g_2^2 g_Y^2 + g_Y^4 - 51g_2^4 - \frac{48g_2^6}{g_Y^2 + g_2^2} \right) \right]. \end{aligned}$$

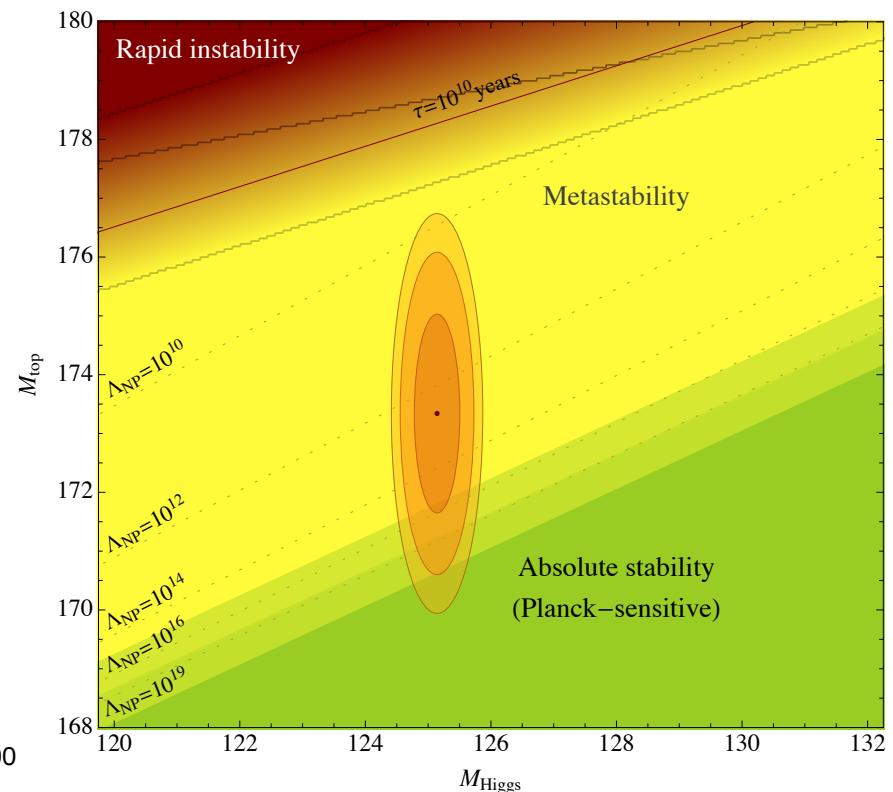
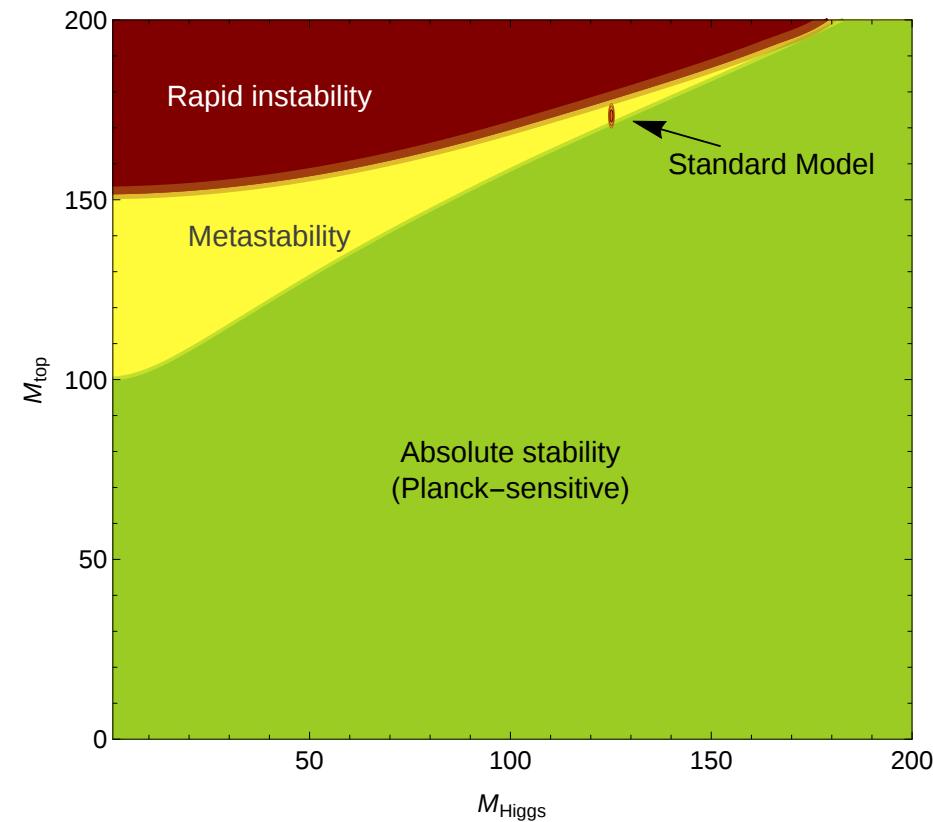
- We don't know the Daisy contribution. But we do know if vanishes in Landau gauge at NLO

$$V^{e^6 \text{daisies}} = \phi^4 \frac{\hbar}{16\pi^2} \left(-\frac{e^2 \lambda \xi}{24} \right) \left[\frac{\hat{\lambda}(\phi)}{\lambda} + \left(1 - \frac{\hat{\lambda}(\phi)}{\lambda} \right) \ln \left(1 - \frac{\hat{\lambda}(\phi)}{\lambda} \right) \right]$$

- Assuming everything works like in scalar QED, we have everything we need for NLO

Results

Absolute stability: for what values of the Higgs and top masses is $V_{\min} = 0$?

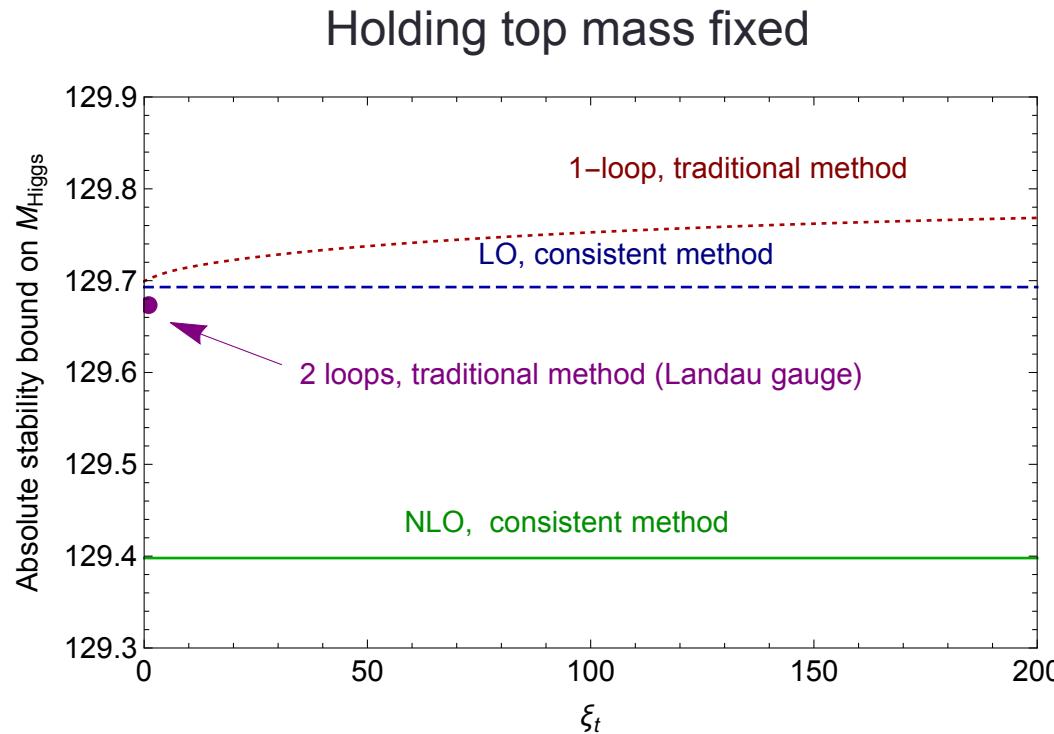


$$m_h^{\text{pole}} = (125.14 \pm 0.23) \text{ GeV}$$

$$m_t^{\text{pole}} = (173.34 \pm 1.12) \text{ GeV}$$

Results

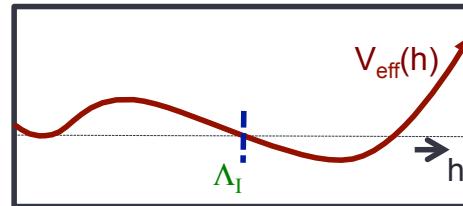
Absolute stability: for what values of the Higgs mass is $V_{\min} = 0$ at fixed top mass?



- Absolute stability bound lowered by 300 MeV
- Larger shift than including the 2-loop V_{eff}

Sensitivity to new physics

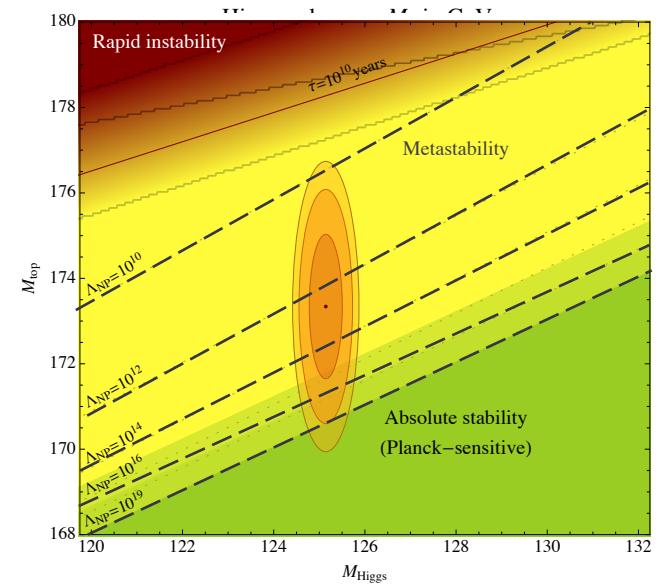
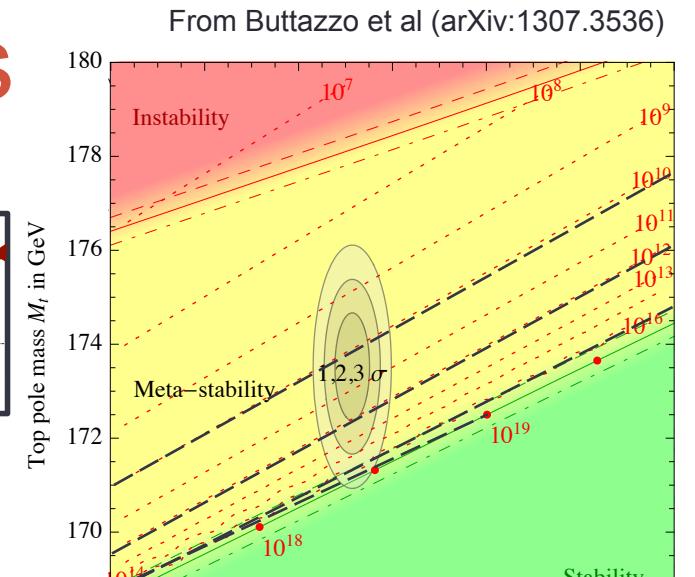
Old way:
when is $\Lambda_I = \Lambda_{NP}$?



- gauge dependent, since Λ_I is gauge-dependent

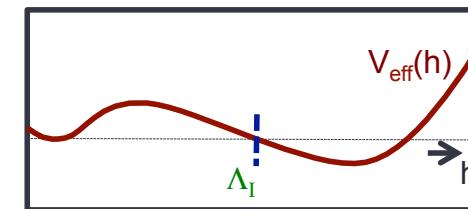
New gauge-invariant way

- Add $\mathcal{O}_6 = \frac{1}{\Lambda_{NP}^2} |H|^6$ to the SM Lagrangian
- See how big Λ_{NP} must be so that $V_{min} = 0$



Planck-sensitivity

Does the tunneling rate depend on quantum gravity?

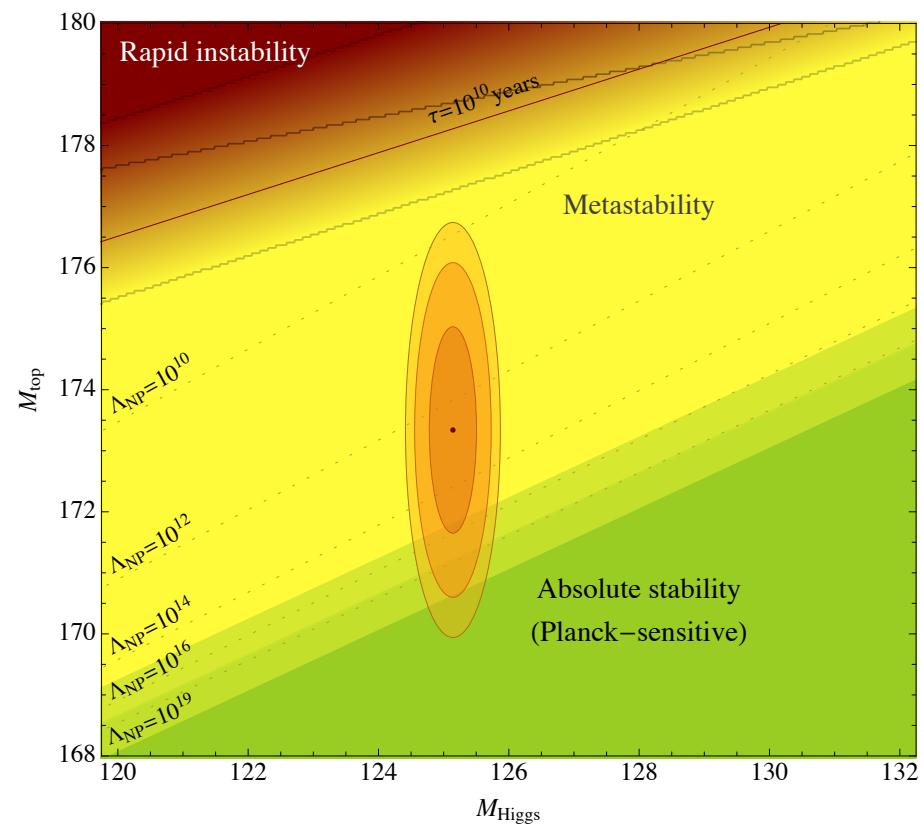


- Guidice, Strumia et al (arXiv:1307.3536):
 - Instability scale below M_{Pl} , so **no**.

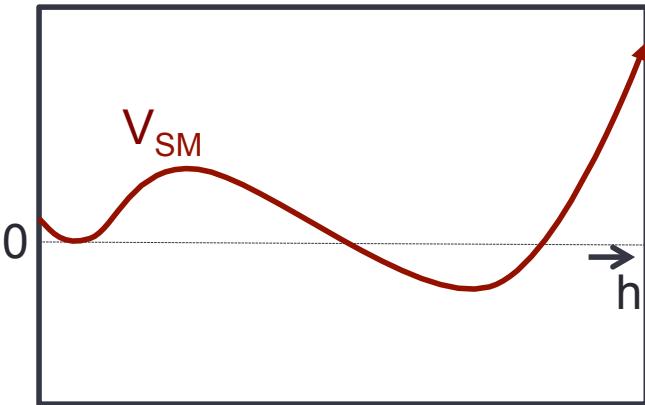
$$\beta_\lambda = 0 \text{ at } \mu = 10^{17} \text{ GeV} < M_{\text{Pl}}$$

- Sher, Brandina et al (arXiv:1408.5302):
 - field at center of bubble is greater than M_{Pl} , so **yes**

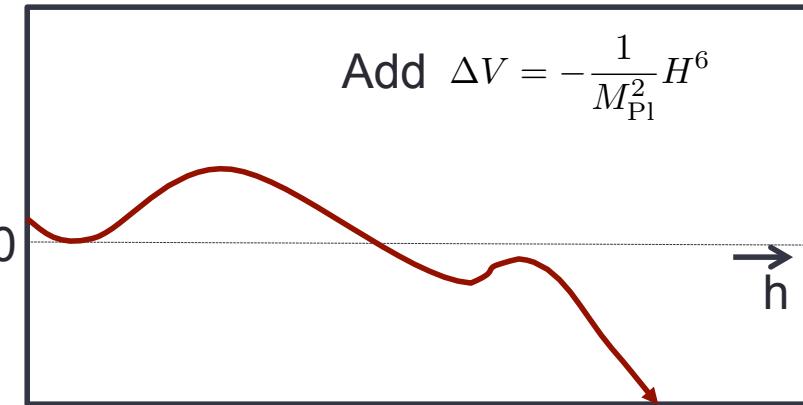
$$\phi_B(r=0) = 10^{19} \text{ GeV} \sim M_{\text{Pl}}$$



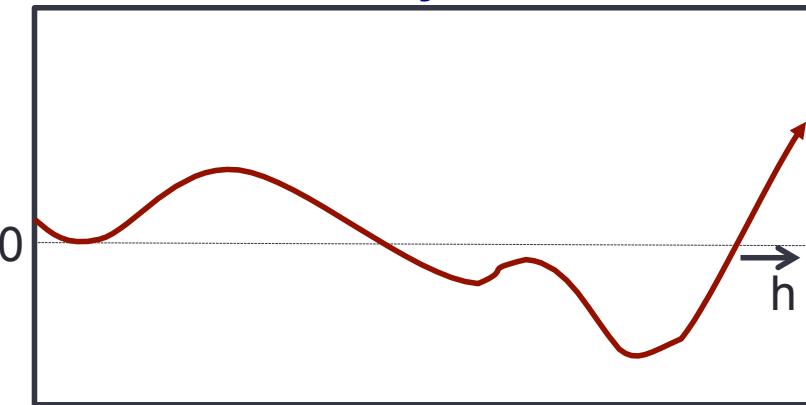
M_{Pl} corrections:



Standard Model potential
Lifetime = 10^{600} years



- **Lifetime = 0 sec**
- Arbitrarily small bubbles form and grow



$$\text{Add } \Delta V = -\alpha \frac{1}{M_{\text{Pl}}^2} H^6 + \beta \frac{1}{M_{\text{Pl}}^2} H^8$$

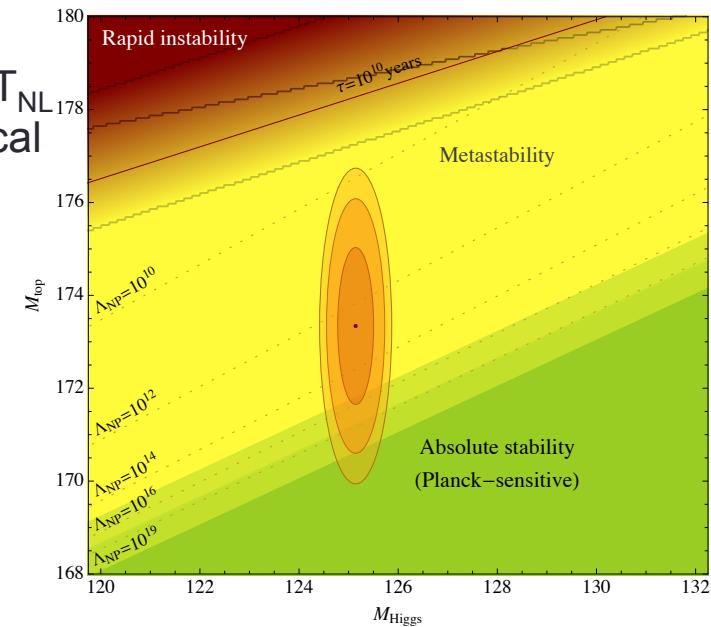
- **Lifetime can be anything!**

- Planck sensitivity not due to coincidence that $\beta_\lambda = 0$ at $\mu \sim M_{\text{Pl}}$
- Tunneling is **non-perturbative** and **always** UV sensitive.

Conclusions

Tunneling involves many **exotic elements** of quantum field theory

- Tunneling rates
 - Two time scales relevant for tunneling: $T_{\text{slosh}} \ll T \ll T_{\text{NL}}$
 - Asymptotic expansions and analytic continuation critical
 - Can be avoided with a direct approach
- Requires consistent use of perturbation theory
 - $\lambda \sim \hbar$ power counting
- UV physics does not decouple
 - Stability is necessarily Planck-sensitive
 - Can make lifetime shorter, not longer



Do we know if the universe is stable?

- Our universe will probably decay, eventually.
- We don't know how long it will last